

Decay Correlations in Pulsed Neutron Beams at ILL, MLZ, and ESS

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PERKEO III

FPSTS19, Oak Ridge

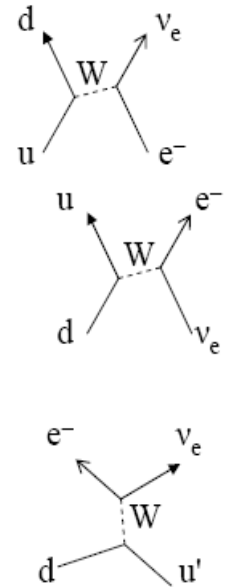
Neutron Decay Data is Useful

Within Standard Model

Many processes share the same Feynman diagram.

Couplings have to be determined in neutron decay!

Primordial element formation (^2H , ^3He , ^4He , ^7Li , ...)	$n + e^+ \rightarrow p + \nu'_e$ $p + e^- \rightarrow n + \nu_e$ $n \rightarrow p + e^- + \nu'_e$
Solar cycle	$p + p \rightarrow ^2\text{H} + e^+ + \nu_e$ $p + p + e^- \rightarrow ^2\text{H} + \nu_e$
Neutron star formation	$p + e^- \rightarrow n + \nu_e$
Pion decay	$\pi^- \rightarrow \pi^0 + e^- + \bar{\nu}_e$
Neutrino detectors	$\bar{\nu}_e + p \rightarrow e^+ + n$
Neutrino forward scattering	$\nu_e + n \rightarrow e^- + p$ etc.
W and Z production	$u' + d \rightarrow W^- \rightarrow e^- + \nu'_e$ etc.



Input parameter to searches for physics **Beyond the Standard Model (BSM)**

Unitarity tests of the first row of the quark mixing CKM matrix on the 10^{-4} level sensitive to new physics at the 10 TeV scale:

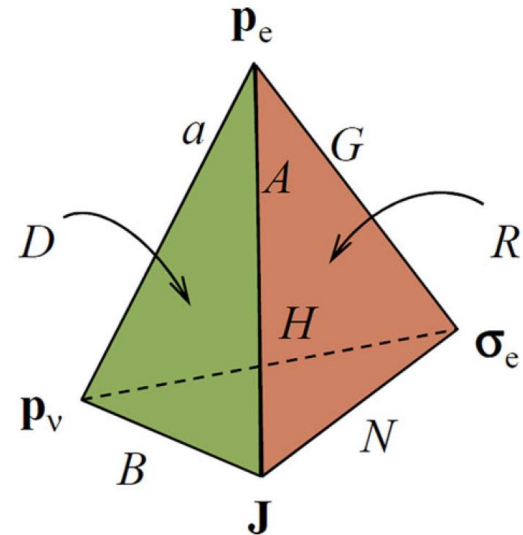
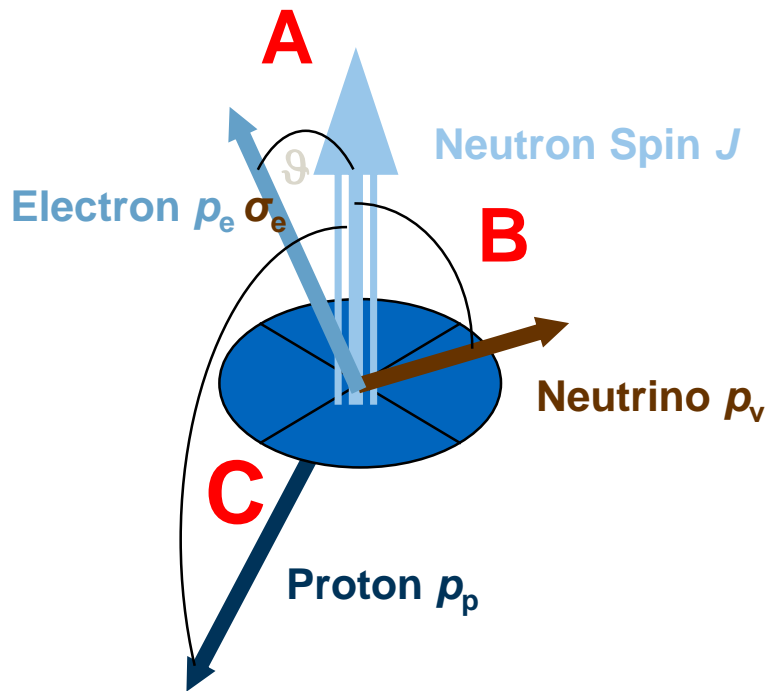
$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 - \Delta$$

Search for new (effective) couplings: scalar, tensor, right-handed

Correlations in Neutron Decay

Determination of $\lambda = g_A/g_V$ from neutron decay via angular correlations:
 (typically) **beta asymmetry A**, or electron-neutrino correlation a

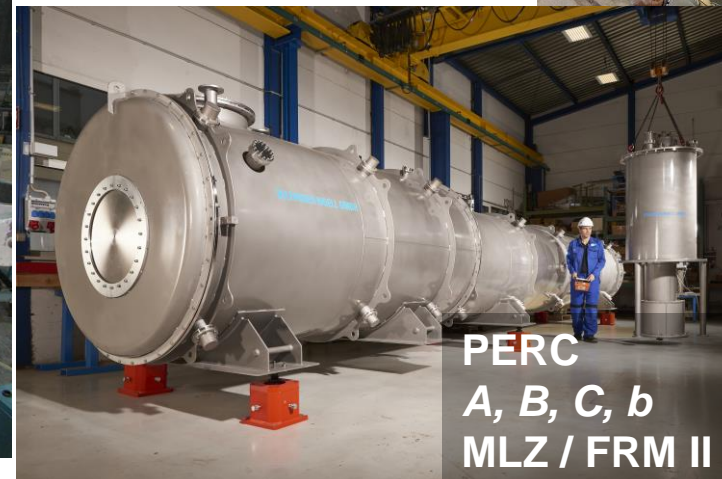
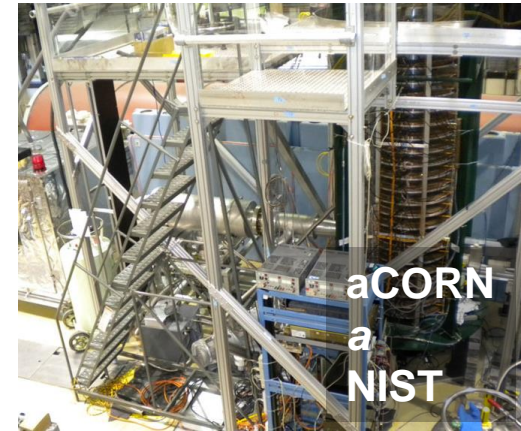
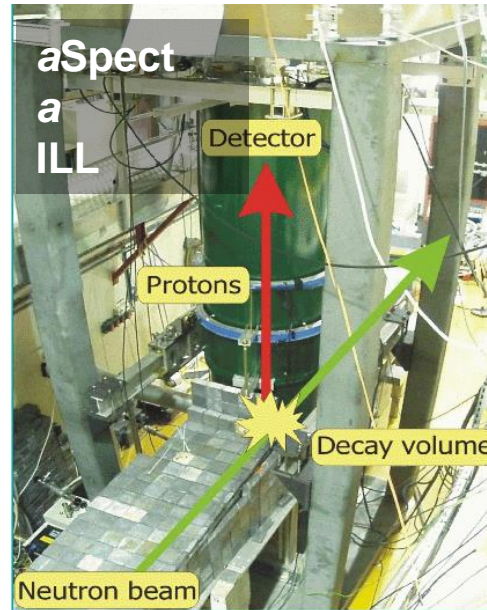
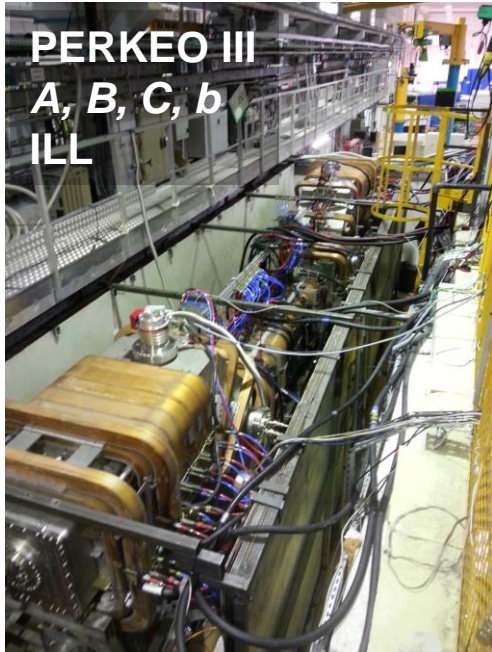
$$A = -2 \frac{\lambda^2 + \lambda}{1 - 3\lambda^2} \quad a = \frac{1 - \lambda^2}{1 - 3\lambda^2}$$



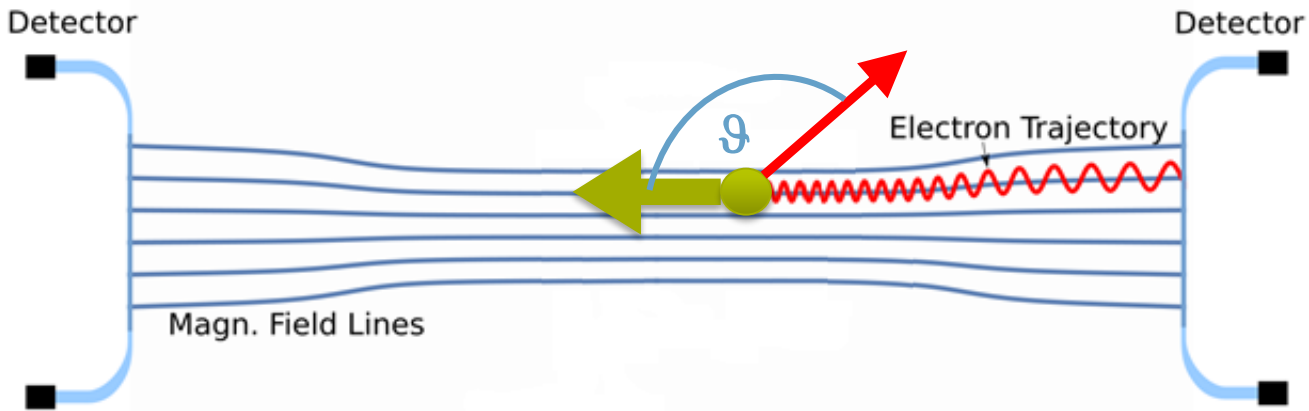
O. Naviliat-Cuncic and M. Gonzalez-Alonso, Ann. Phys. 525, 8–9, 600–619 (2013)
Dubbers and Schmidt, Rev. Mod. Phys (2012)

Typically, **specialised** instruments / setups required for different observables.

Current Neutron Correlation Experiments



PERKEO: Measuring Beta Asymmetry



Electron angular distribution:

$$W(\vartheta, E) = 1 + \frac{v}{c} A \cos \vartheta$$

Within Standard Model:

$$A = -2 \frac{\lambda^2 + \lambda}{1 + 3\lambda^2} \quad \lambda = \frac{g_A}{g_V}$$

Magnetic field for spin alignment

Integration over hemispheres:
 $2 \times 2 \pi$ detection

$$\cos \vartheta \rightarrow \frac{1}{2}$$

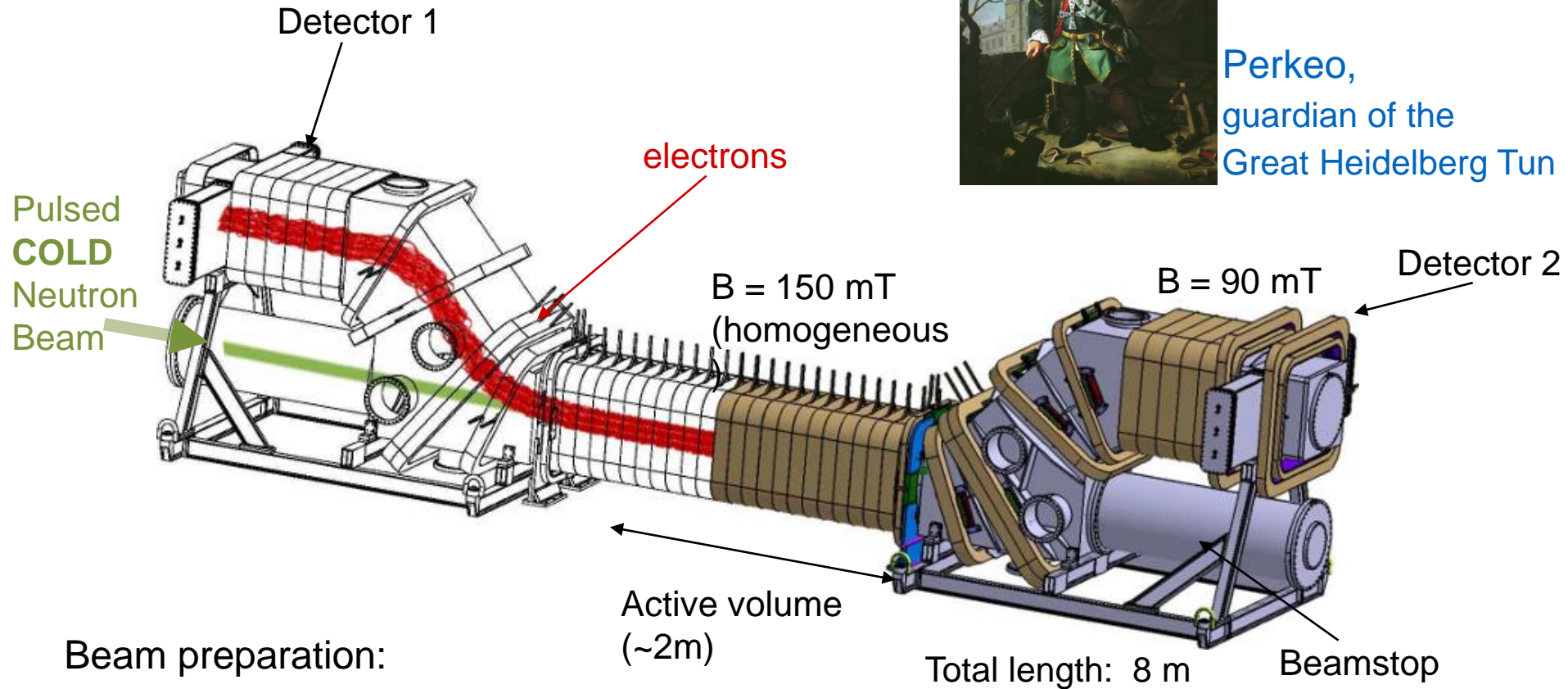
Experimental asymmetry, polarisation P

$$A_{\text{exp}} = \frac{N^{\uparrow} - N^{\downarrow}}{N^{\uparrow} + N^{\downarrow}} = \frac{1}{2} \frac{v}{c} P A$$

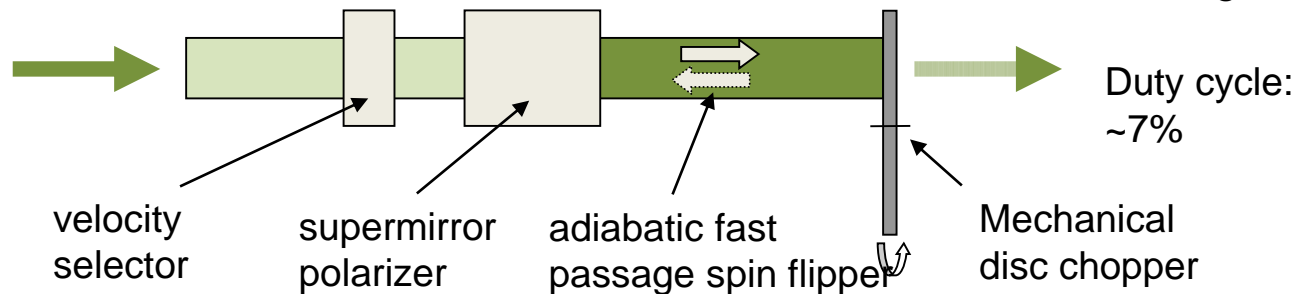
Spectrometer PERKEO III



Perkeo,
guardian of the
Great Heidelberg Tun



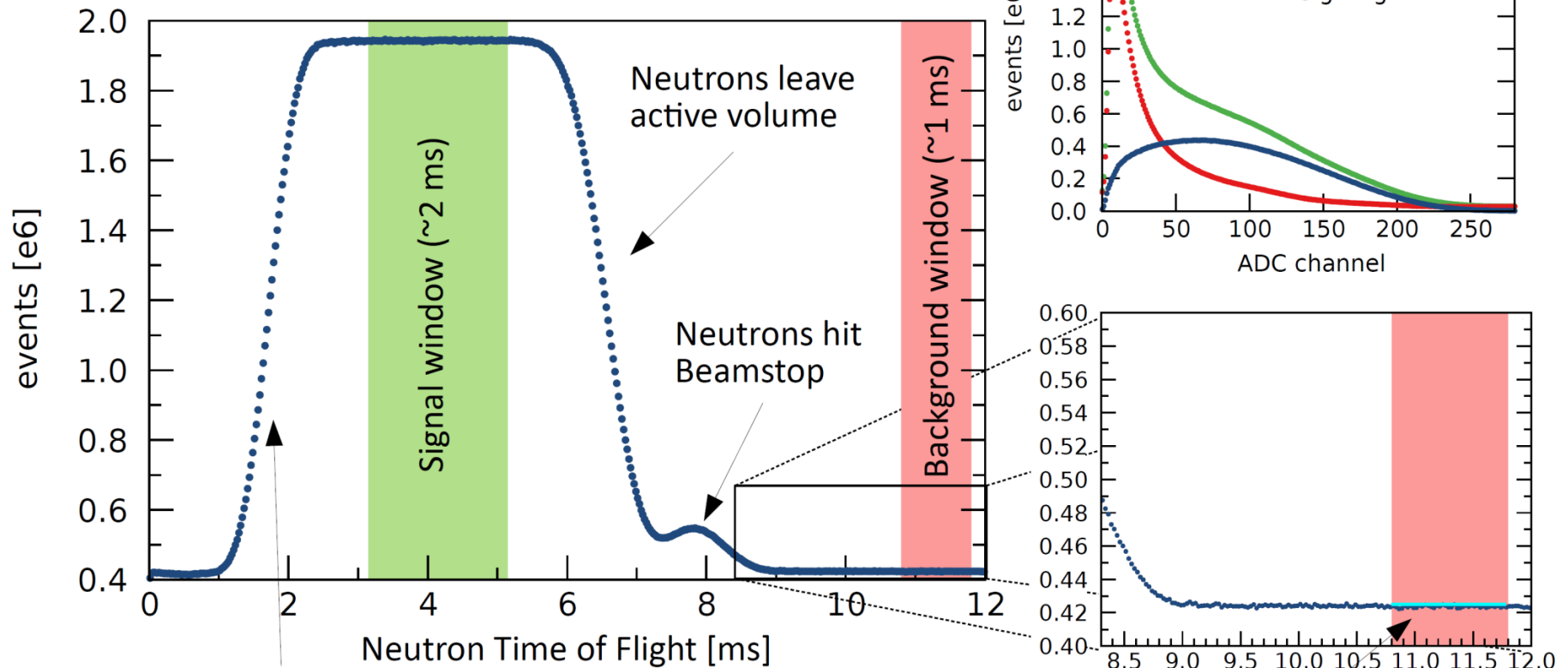
Beam preparation:



~ 50.000 decays / sec
in continuous beam
time avg. $\sim 200 \text{ s}^{-1}$ in
pulsed mode

PERKEO III: Pulsed Neutron Beam

Energy range
 $300 \text{ keV} < E < 700 \text{ keV}$



Neutrons enter active volume

Related Uncertainties:

Time dependence $\Delta A/A = 0.8 \times 10^{-4}$

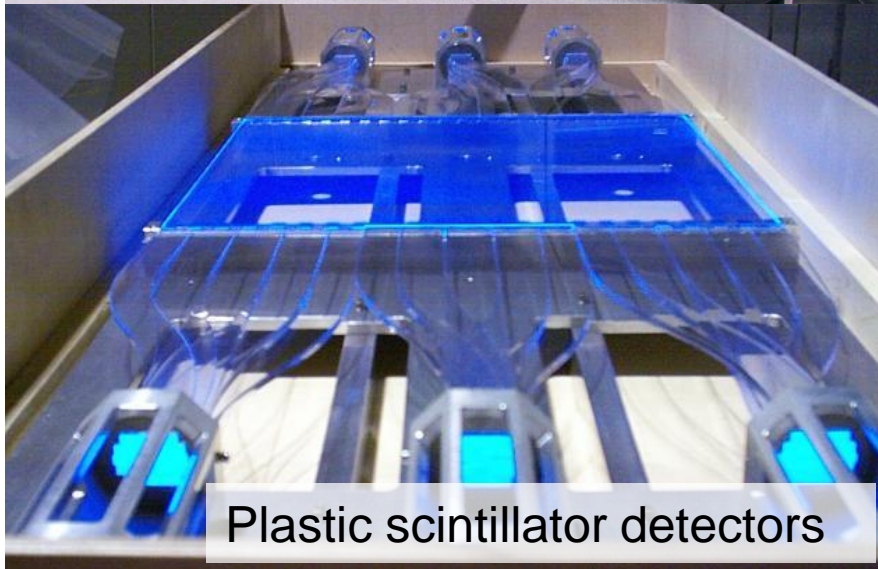
Chopper disc uniformity $\Delta A/A = 0.7 \times 10^{-4}$

Slope is zero compatible on 10^{-4} level

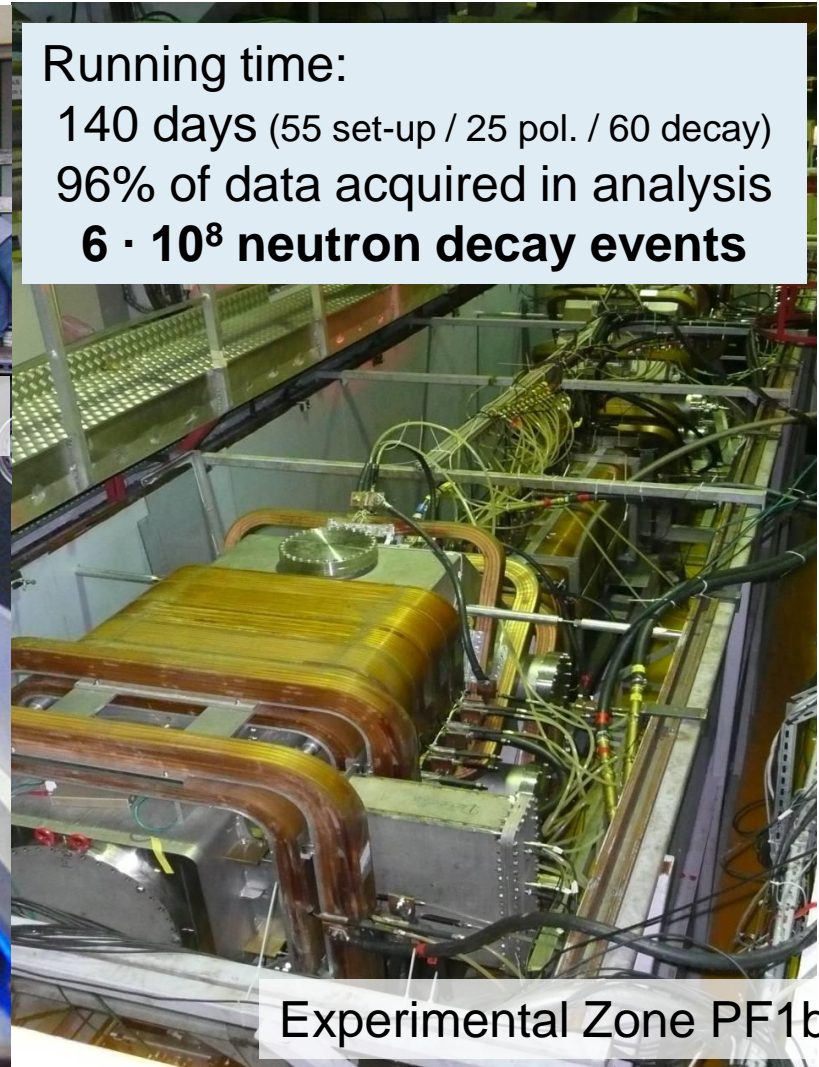
Temporary Installation at PF1b, ILL



One of two trucks



Plastic scintillator detectors



Experimental Zone PF1b

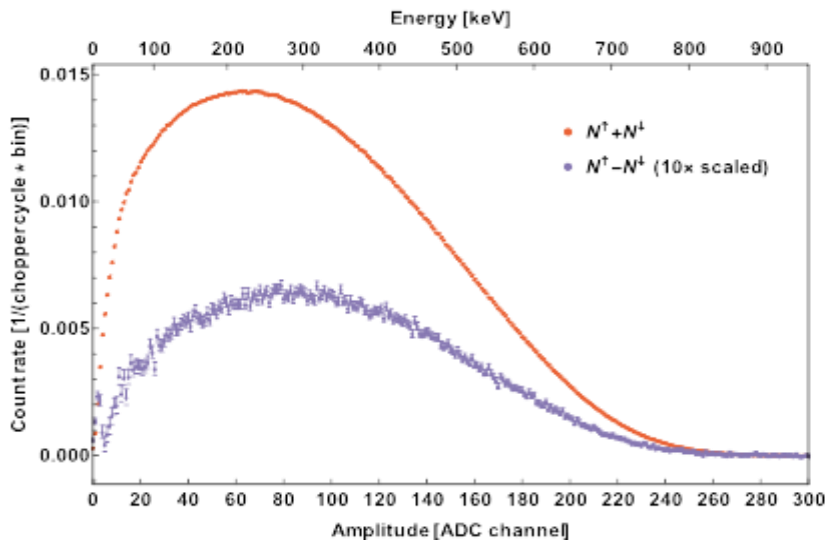
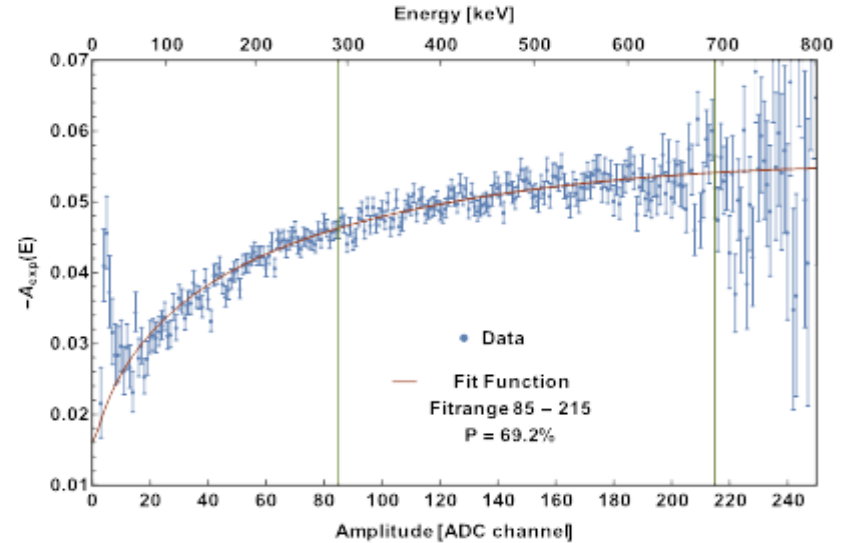
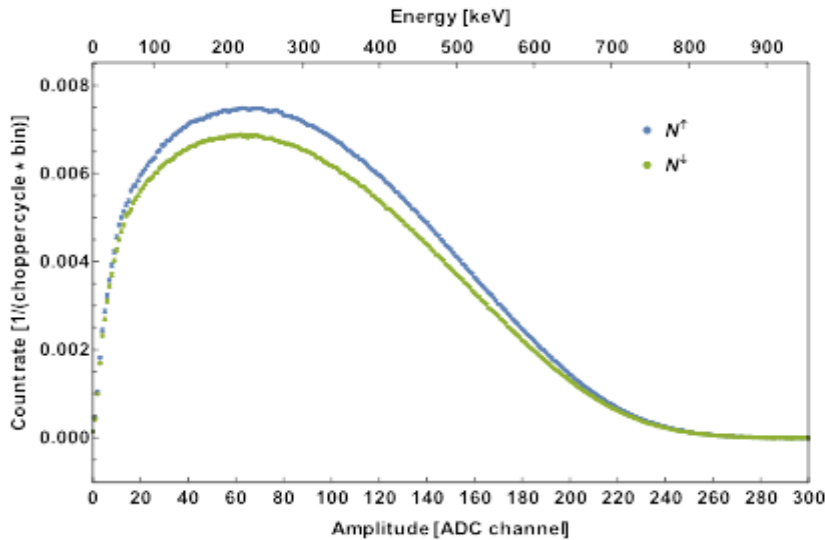
Running time:

140 days (55 set-up / 25 pol. / 60 decay)

96% of data acquired in analysis

$6 \cdot 10^8$ neutron decay events

Asymmetry Extraction



Fit a single free parameter λ to experimental asym.

$$A_{exp}(E_e) = \frac{N^\uparrow(E_e) - N^\downarrow(E_e)}{N^\uparrow(E_e) + N^\downarrow(E_e)} = \frac{1}{2} P_n \frac{v}{c} A$$

Largest neutron decay data set

6×10^8 events in analysis (1 of 4 subsets shown)

Statistical Uncertainty: $\Delta A/A = 14 \times 10^{-4}$

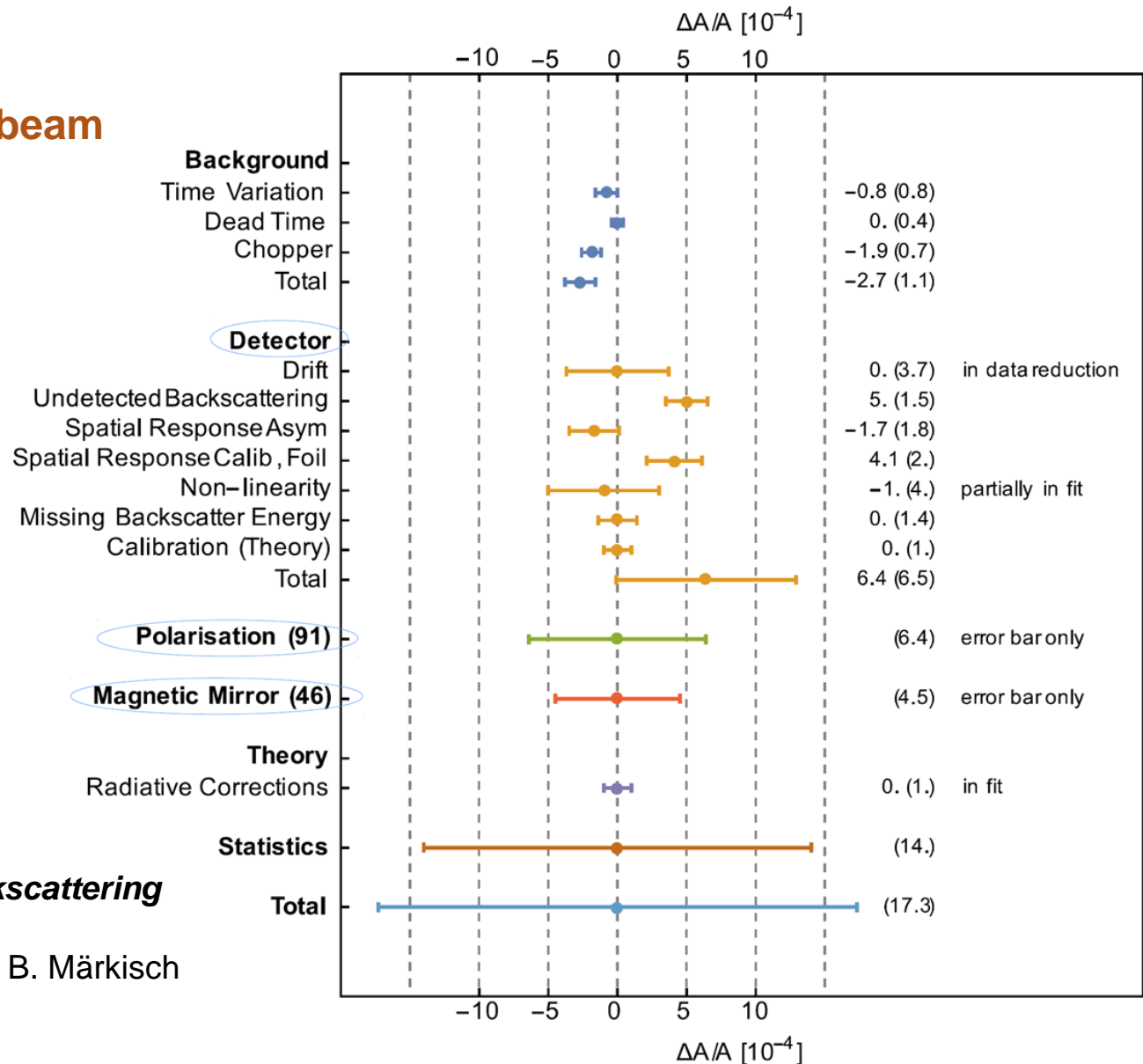
Summary of Corrections and Uncertainties

First measurement of λ using a pulsed beam

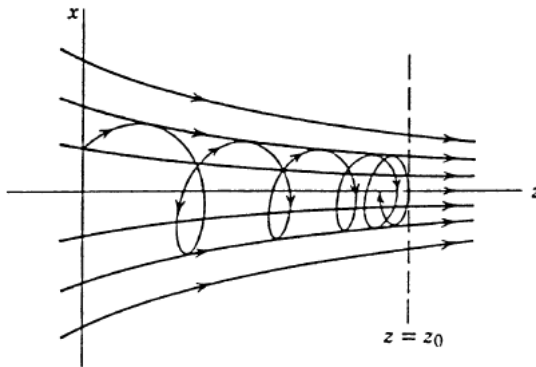
see also

Undetected Electron Backscattering in PERKEO III

C. Roick, H. Saul, H. Abele, B. Märkisch
arXiv:1905.10189



Magnetic Mirror Effect



Magnetic flux through cross section of gyration is *adiabatic invariant*

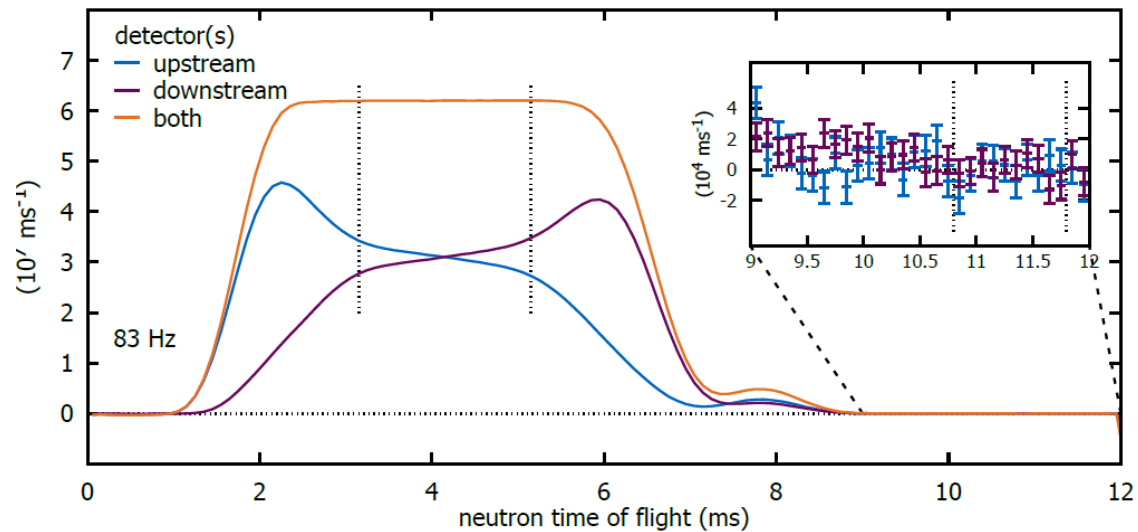
$$B_0 \times r_0^2 = B_1 \times r_1^2$$

Critical angle for reflection

$$\Theta_c = \arcsin \sqrt{\frac{B_1}{B_0}}$$

Magnetic field drops by 1% towards the ends of the decay volume:

Significant change in on solid angle and rate on **single** detector.

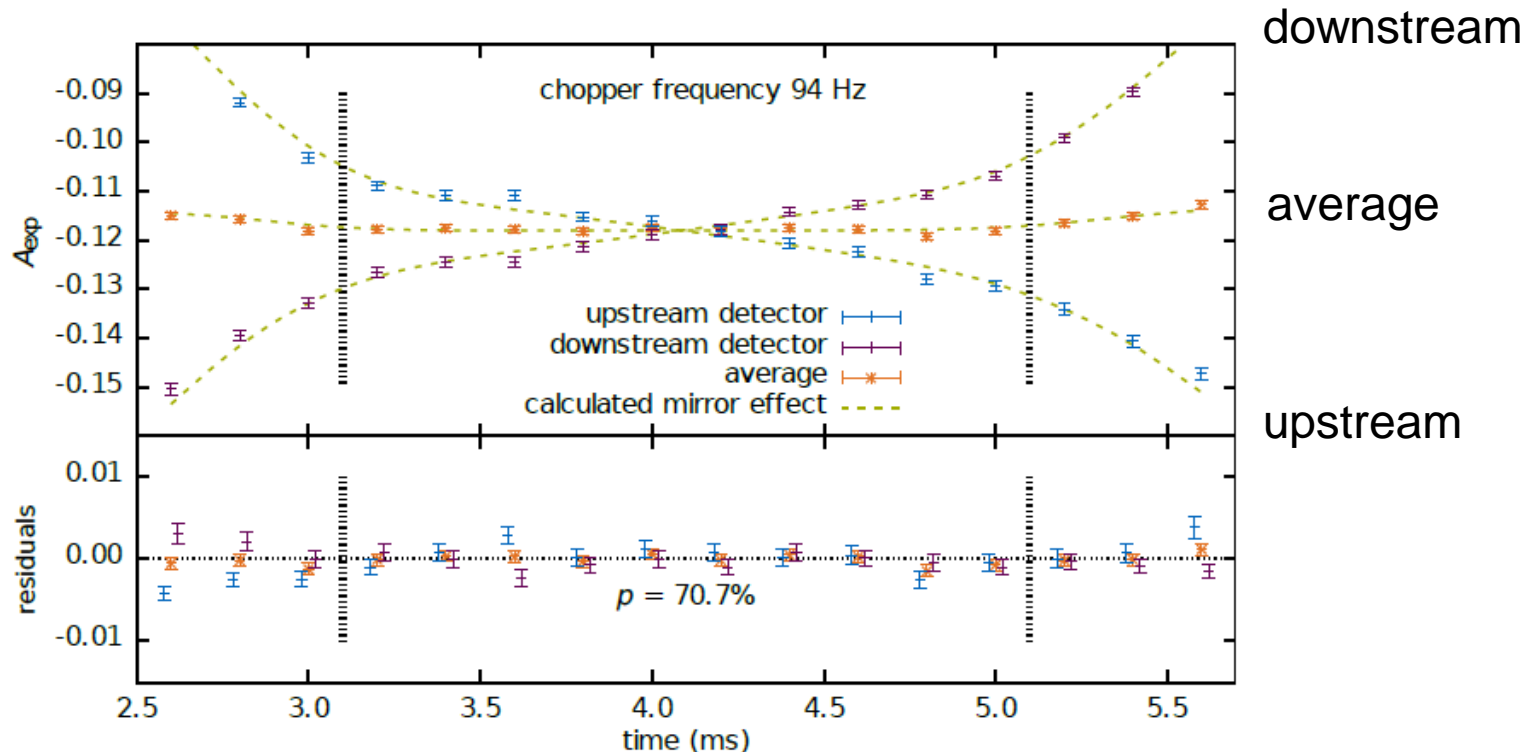


Magnetic Mirror Effect

Most of the effect cancels by **averaging** detectors.

Remaining **correction** calculated from measurements of the magnetic field and neutron pulse. Correction: $\Delta A/A = 46.1(4.5) \times 10^{-4}$

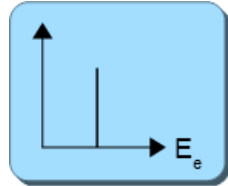
Selection of time-window during analysis.



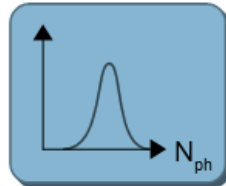
Detector Model

Major improvements to precise description of the detector response enable consistent energy-dependent analysis (H. Saul, C. Roick).

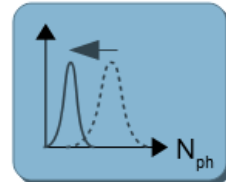
Free parameters: Non-linearity, gain, photo-electrons, norms



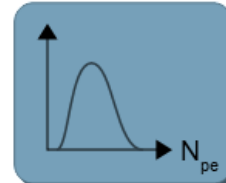
Electrons:
discrete energy
or spectrum



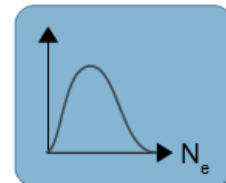
Scintillation:
 $N_{ph} = f(E_e)$
Poisson statistics



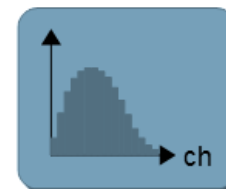
Photon transport:
 $N'_{ph} = f(E_e, x, y)$
binomial statistics



Photon to photoelectron conversion:
 $N_{pe} = f(N_{ph})$
binomial statistics



Electron multiplication (PMT):
 $N_e = f(N_{pe})$
Poisson statistics at N=19 stages



Signal processing + charge integration:
 $A_{QDC} = f(N_e)$
Gaussian noise

Non-linearity of
scintillation
light production

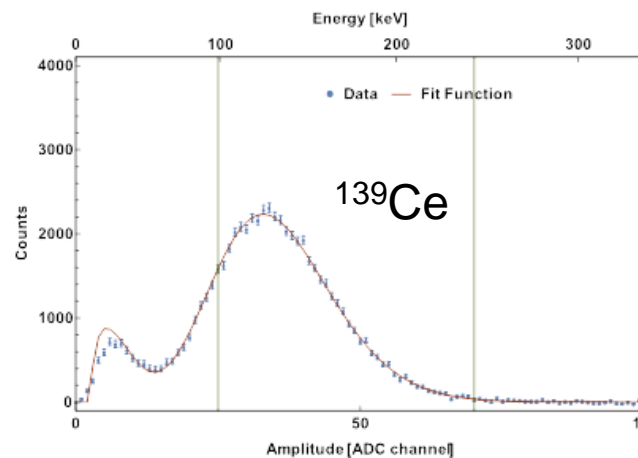
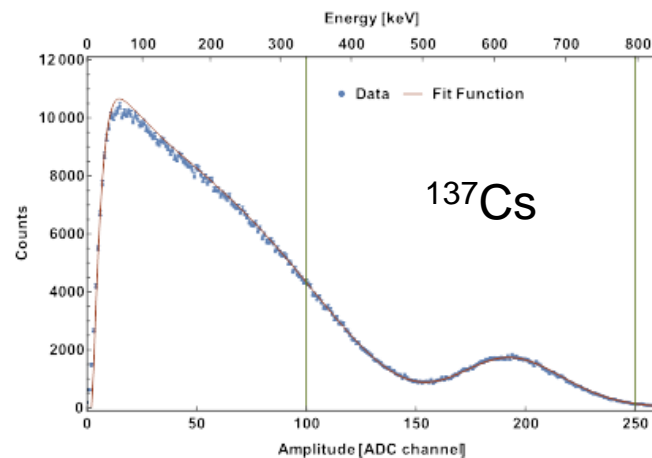
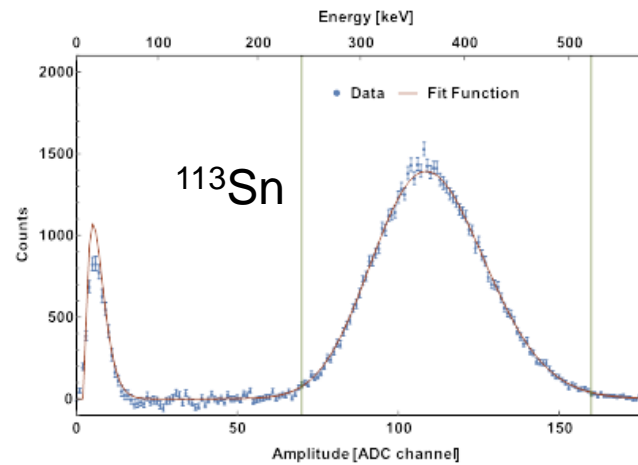
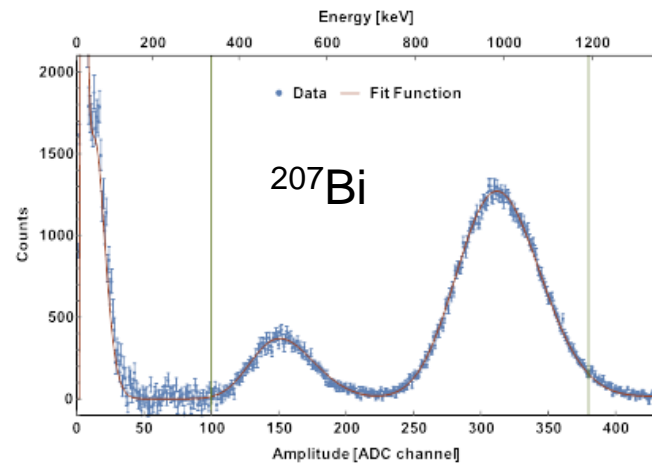
Non-uniformity of
detector response

Higher moments of
the distribution

Non-linearity of
electronics

PERKEO III Detector Calibration Fit

Calibration, drift monitoring and uniformity scans using electron-conversion sources



114 full calibration sets measured in ~60 days

Simultaneous fit:
 $\chi^2/\text{NDF} = 1.0 - 1.3$

Free parameters:
Non-linearity, gain, photo-electrons, norms

Related Uncertainties

Sources: $\Delta A/A = 1 \times 10^{-4}$

Statistics: $\Delta A/A = 0.1 \times 10^{-4}$

Non-linearity: $\Delta A/A = 4 \times 10^{-4}$

Stability: $\Delta A/A = 3.7 \times 10^{-4}$

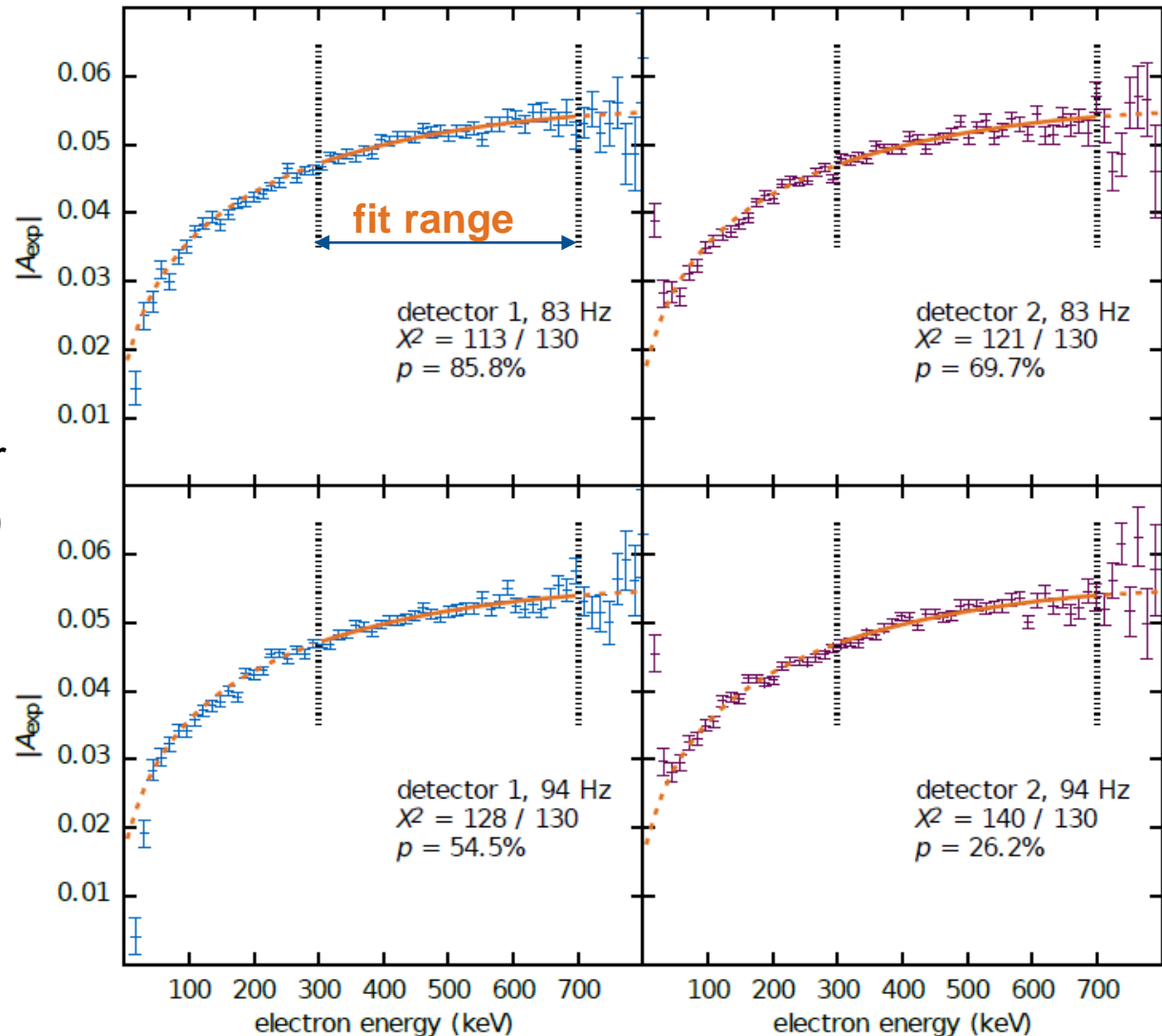
Asymmetry: Four datasets

Two chopper frequencies (background systematics), two detectors

Fit to energy-dependence of experimental asymmetry $A_{\text{exp}}(E_e)$,

Model includes detector response (no unfolding)

Only a single free parameter: λ



PERKEO III Result

Analysis blinded by separate analysis by independent teams:

- electron measurement,
- neutron polarisation: opaque ^3He spin filters,
- magnetic mirror effect

$$A = -2 \frac{\lambda^2 + \lambda}{1 + 3\lambda^2}$$

$$\lambda = -1.27641(45)_{\text{stat}}(33)_{\text{sys}}$$

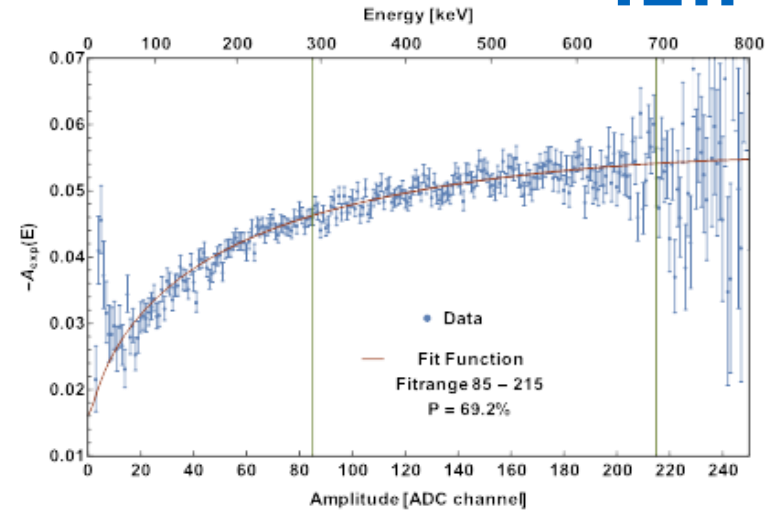
$$= -1.27641(56)$$

$$A = -0.11985(17)_{\text{stat}}(12)_{\text{sys}}$$

$$= -0.11985(21).$$

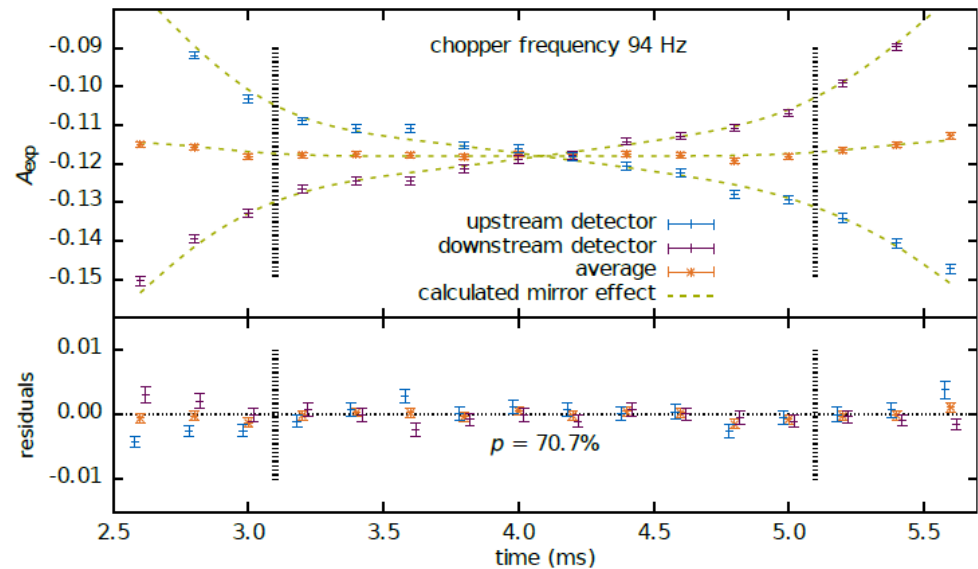
$$\frac{\Delta\lambda}{\lambda} = 4.4 \times 10^{-4}$$

B. Märkisch, H. Mest, H. Saul, X. Wang, H. Abele, D. Dubbers, M. Klopff, A. Petoukhov, C. Roick, T. Soldner, D. Werder, Phys. Rev. Lett. 122, 222503 (2019)



Time-dependent mirror effect

Time-avg. correction: $\Delta A/A = 46.1(4.5) \times 10^{-4}$



Nucleon Axial Coupling: Status

All new results **consistent** – but disagree with older measurements.
 Newer measurements of A have order of magnitude **smaller corrections**.
 UCNA and PERKEO III: **blinded analysis**.

PERKEO III:

$$\lambda = -1.27641(56),$$

$$\frac{\Delta\lambda}{\lambda} = 4.4 \times 10^{-4}$$

PDG 2018:

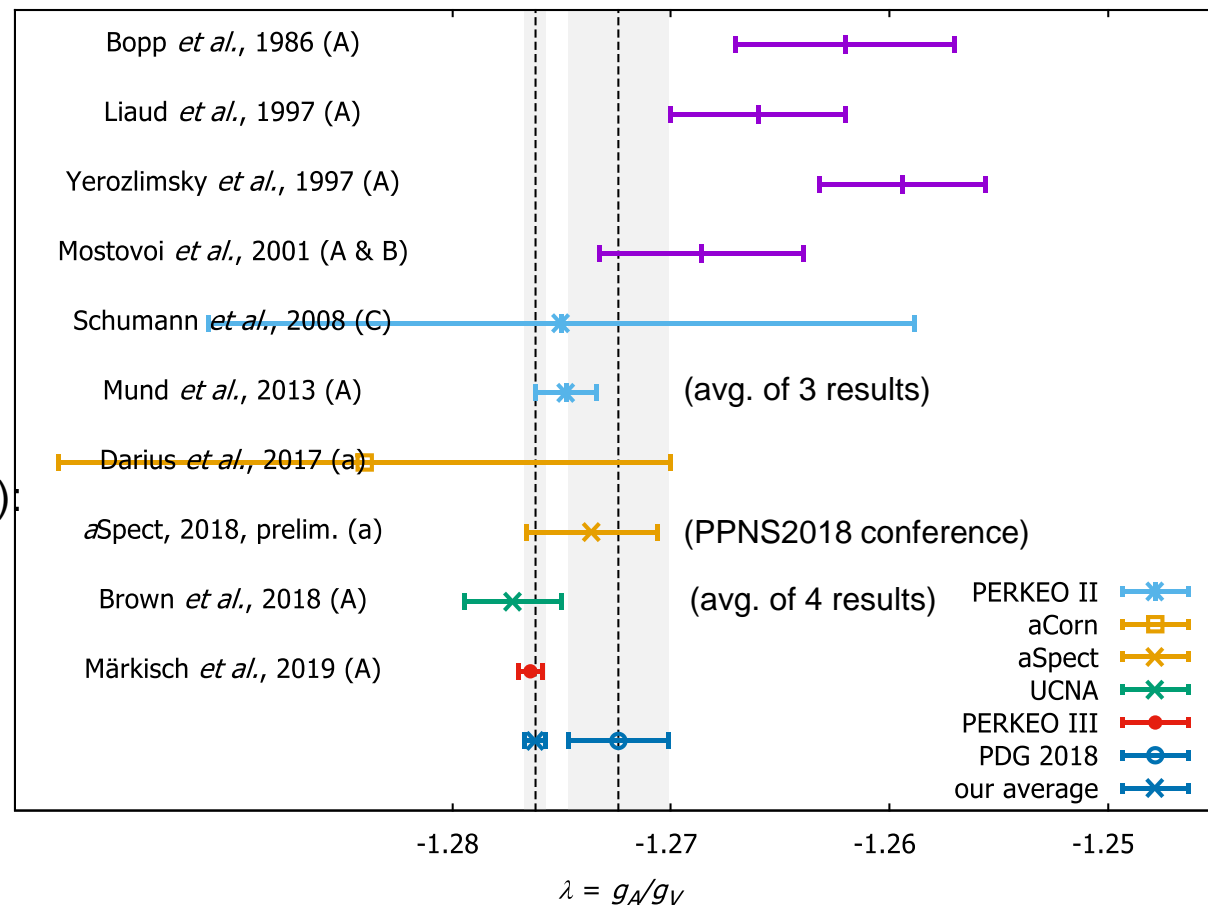
$$\lambda = -1.2724(23); S = 2.4$$

New average (all measurements):

$$\lambda = -1.2756(10); S = 2.15$$

Only UCNA, PERKEO II & III:

$$\lambda = -1.2762(5)$$

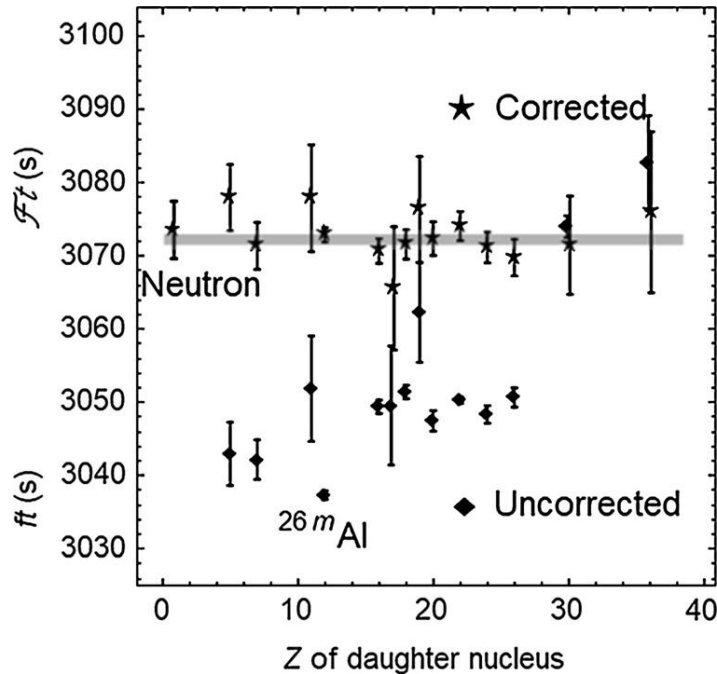


Comparison to Superallowed Nuclear Decays

Neutron data confirms nuclear data with competitive precision –
without nuclear structure effects !

Vector part of neutron Ft

$$Ft_{nV} \equiv f t_{nV} (1 + \delta'_R) = \frac{1}{2} \ln 2 f \tau_n (1 + 3\lambda^2) (1 + \delta'_R)$$



Using updated (conservative) *world averages*:

$$\lambda = -1.2756(10); S = 2.15$$

$$\tau = 879.7(8); S = 1.9$$

Nuclear data from Hardy & Towner, PRC 91 (2015) 025501. D. Dubbers, et al., Phys. Lett. B 791, 6-10 (2019)

PERKEO III result only:

$$\begin{aligned} V_{ud} &= \left(\frac{5099.34 \text{ s}}{\tau_n (1 + 3\lambda^2) (1 + \Delta_R)} \right)^{1/2} \\ &= 0.97301(10)_{RC(44)} \tau_n(35) \lambda \\ &= 0.97301(58), \end{aligned}$$

Superallowed decays:

$$V_{ud} = 0.97395(23)$$

C.-Y. Seng, M. Gorchtein, M Ramsey-Musolf
Phys. Rev. D 100 (2019)

No Dark Side to Neutron Decay!

Assuming V-A, neutron lifetime can be inferred using our (conservative) world-average of neutron decay data, *including all measurements*.

$$\tau_{\beta}^{\lambda} = \frac{2}{\ln 2} \frac{\overline{F}t_{0^{+} \rightarrow 0^{+}}}{f(1 + \delta'_{R})(1 + 3\lambda^2)} = \frac{5172.3(1.1) \text{ s}}{1 + 3\lambda^2}$$

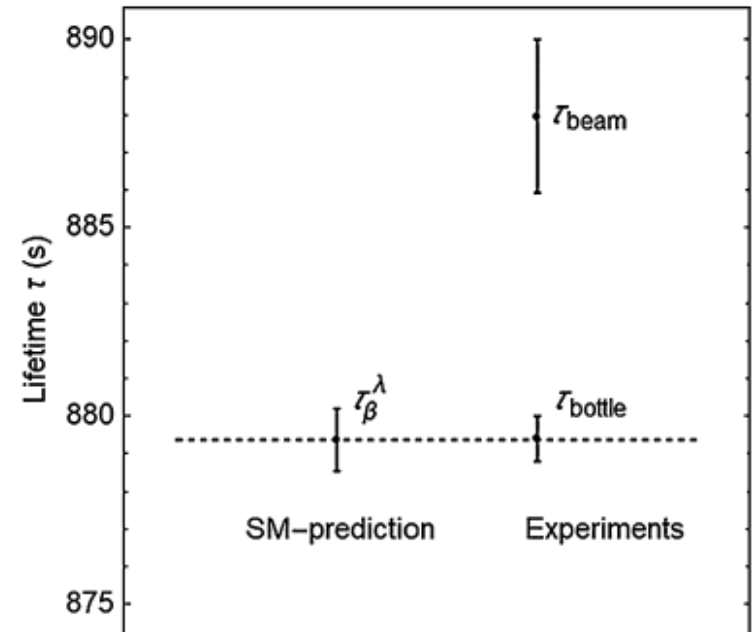
Globally, there's very little room for new physics: $BR_{DM} < 0.30\%$ 95% C.L.

Exotic decay channels are not the cause of the neutron lifetime anomaly.

D. Dubbers, H. Saul, B. Märkisch, T. Soldner and H. Abele Phys. Lett. B 791, 6-10 (2019)

Some channels excluded by experiments (UCN τ , UCNA, Perkeo II), contradiction to the existence of heavy neutron stars, ...

See also: A. Czarnecki, W.J. Marciano, A. Sirlin, Phys. Rev. Lett. 120 (2018) and arXiv:1907.06737



The next generation: PERC (Proton Electron Radiation Channel) at MLZ

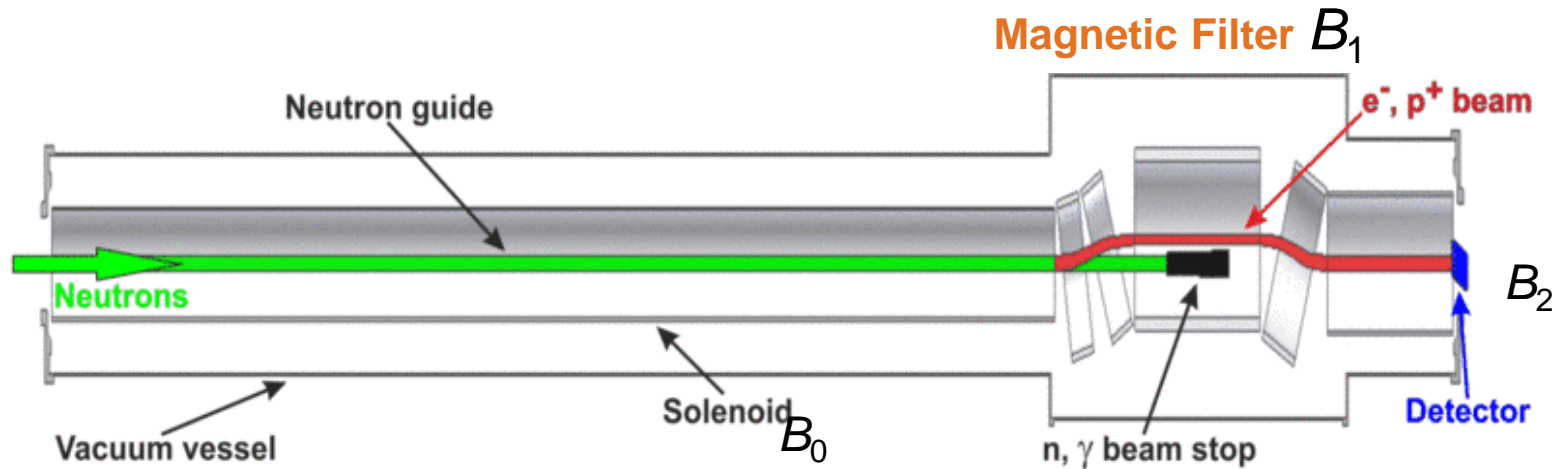
Goal: Order of magnitude improvement. New observables.



Priority Programme SPP1491 of
the German Research Foundation



Proton Electron Radiation Channel (PERC)



Active volume in a 8 m long *neutron-guide*, $B_0 = 1.5$ T:
phase space density and statistics

Magnetic Filter, $B_1 = 6$ T: phase space, systematics
(solid angle, backscatter suppression) $B_1/B_0 = 2 \dots 12$

Source for specialised spectrometers

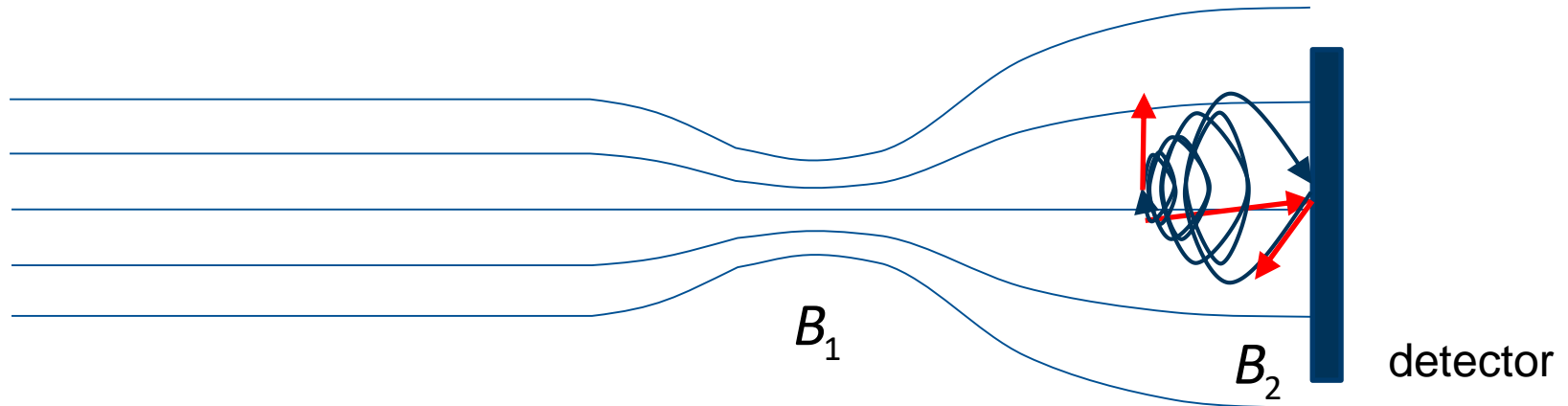
D. Dubbers *et al.*, *Nucl. Instr. Meth. A* **596** (2008) 238 and arXiv:0709.4440

Design of the Magnetic System of the Neutron Decay Facility PERC

X. Wang, C. Drescher *et al.* (PERC Collaboration), EPJ Web Conf., in press, arxiv:1905.10249

Backscatter Suppression in PERC

Backscattering typically major source of systematic error in electron spectroscopy, but:

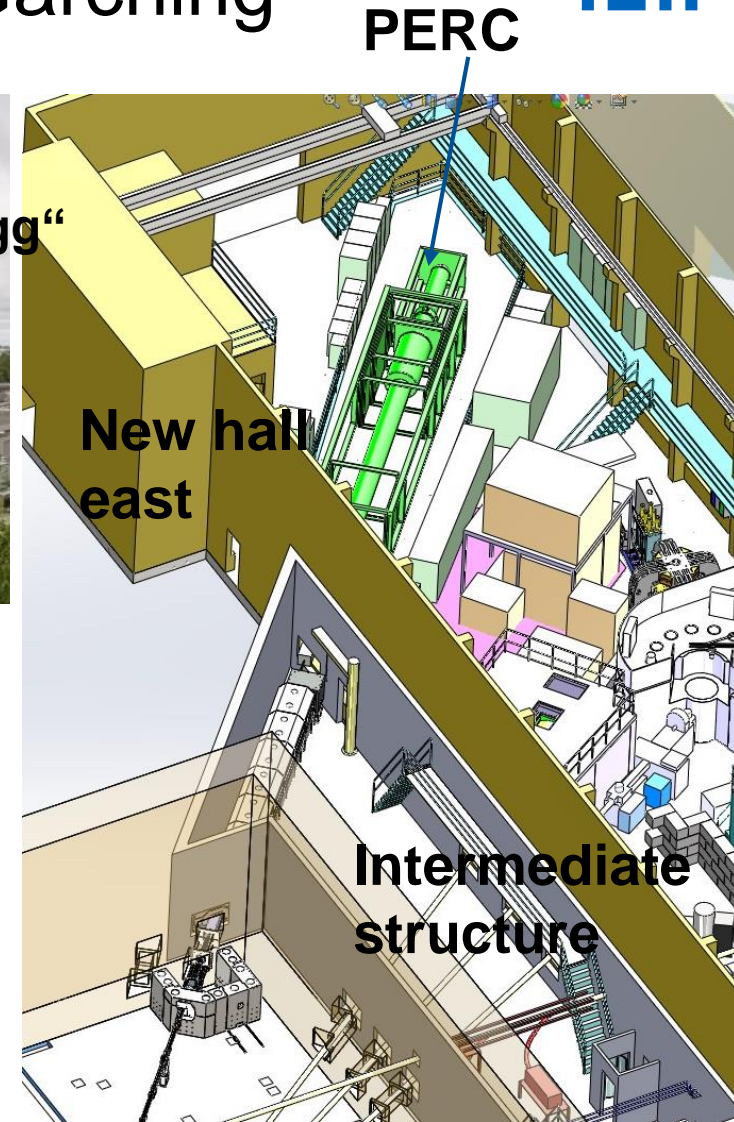
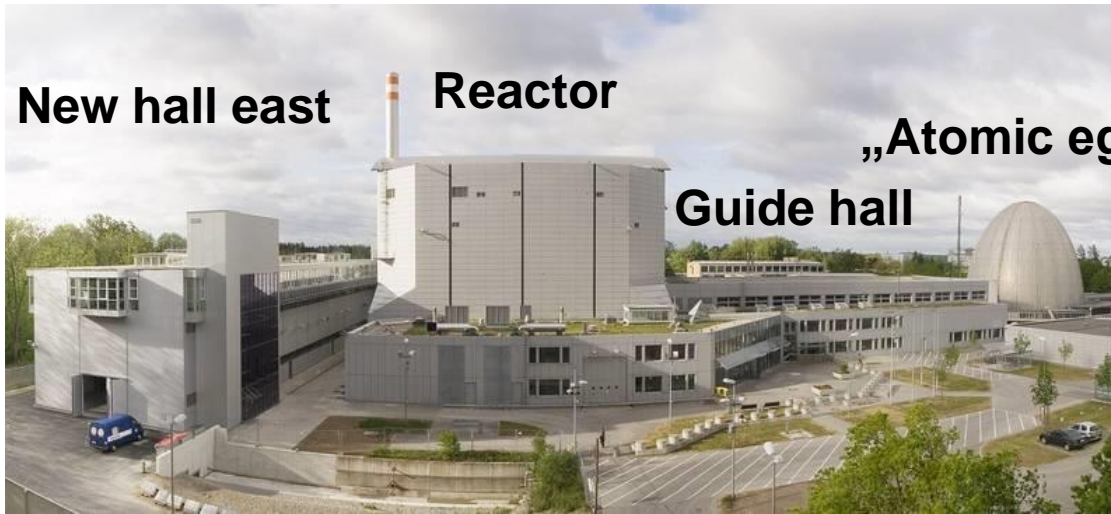


Incident angle on detector $\theta_2 < 15^\circ$ \rightarrow backscattering **small**
Efficient solid angle **suppression** by magnetic mirror

$$\frac{\Omega}{2\pi} = \frac{B_2}{2B_1} \approx 1/32$$

\rightarrow One detector only is sufficient. But PERC will still have an upstream backscatter detector detector with $\sim 95\%$ beam coverage: active electron “dump”

Beam Site Mephisto, FRM II, Garching



Neutron guide:
length 40 m, $R = 3000$ m, $m = 2.5$
Expected intensity equal to PF1b at ILL,
 $2 \times 10^{10} \text{ s}^{-1} \text{ cm}^{-2}$
Only very few neighbours:
low ambient background

Experimental area: 5m x 25m

Reactor core

Beamline Shielding for MEPHISTO



**Guides, selector, ...
READY to be installed.**

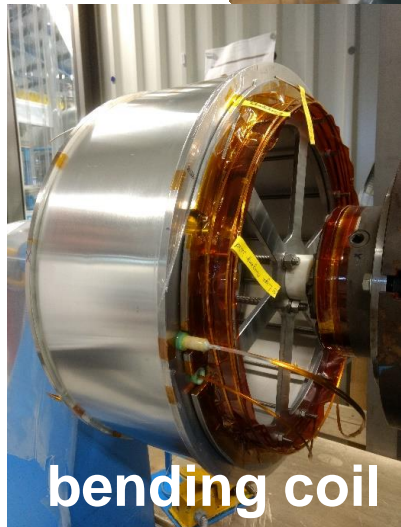
PERC Magnet Status

Delivery Spring 2020



6T filter section

(Needs to be redone - ongoing)



bending coil



8m solenoid



space frame

Cold Beamline for Particle Physics at ESS

ESS design goal is same time average neutron flux as ILL.

Peak brightness in pulse: $30 \times$ ILL

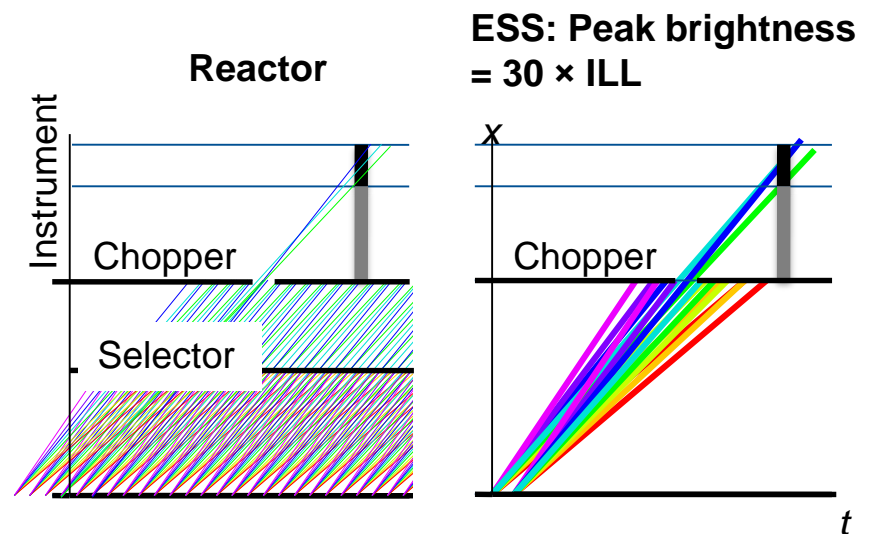
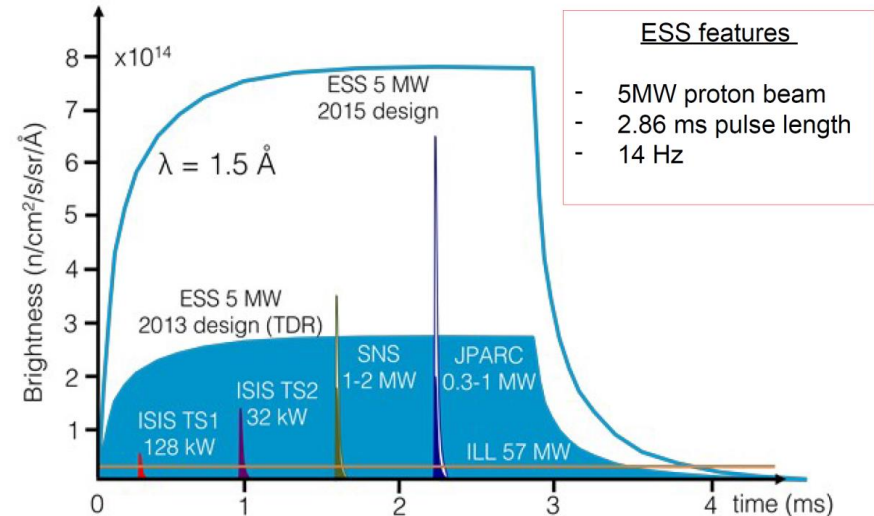
Using pulsed beam for particle physics already at reactor sources!

Statistics gain factor for a PERC-like system: $\times 15$!

Status: particle physics ranked first on “ESS instrument suite – Capability Gap Analysis”

ANNI – A pulsed cold neutron beam facility for particle physics at the ESS

T. Soldner, H. Abele, G. Konrad, B. Märkisch, F. Piegsa, U. Schmidt, C. Theroine, P. Torres Sánchez, EPJ Web Conf. (2019), in press, arXiv:1811.11692



Proposed Magnet Concept for ANNI / ESS

(Nordita Workshop, Stockholm, Dec 2018)

Long decay volume

Either with guide and longer - like PERC,
or with larger cross section and shorter for divergence acceptance - like PERKEO III

Beam separation

enables continuous polarimetry - but need shielding!

Upstream detector

Fully capable upstream detector, well shielded against beam background:

Backscatter detection. Veto via time-of-flight, see Roick et al., PRC 2018

Ensure full energy reconstruction for primary detector.

Magnetic barrier field for downstream detector

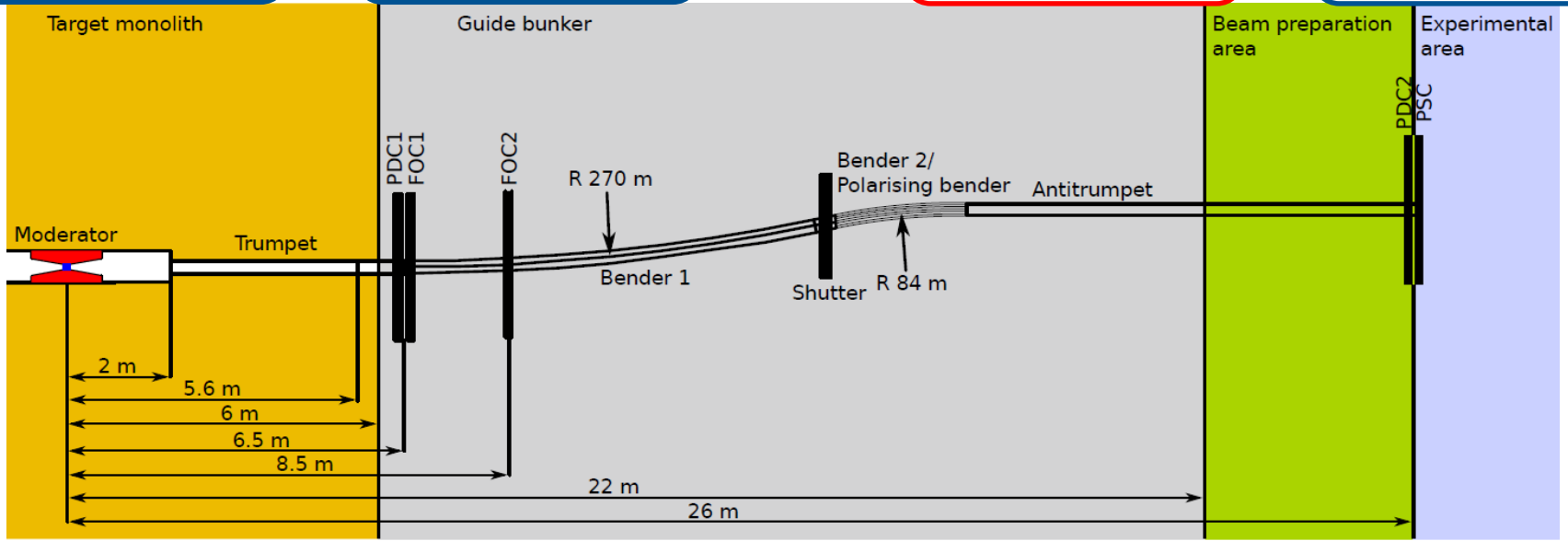
Improved systematics

Enable slight gradient field in decay volume for improved extraction



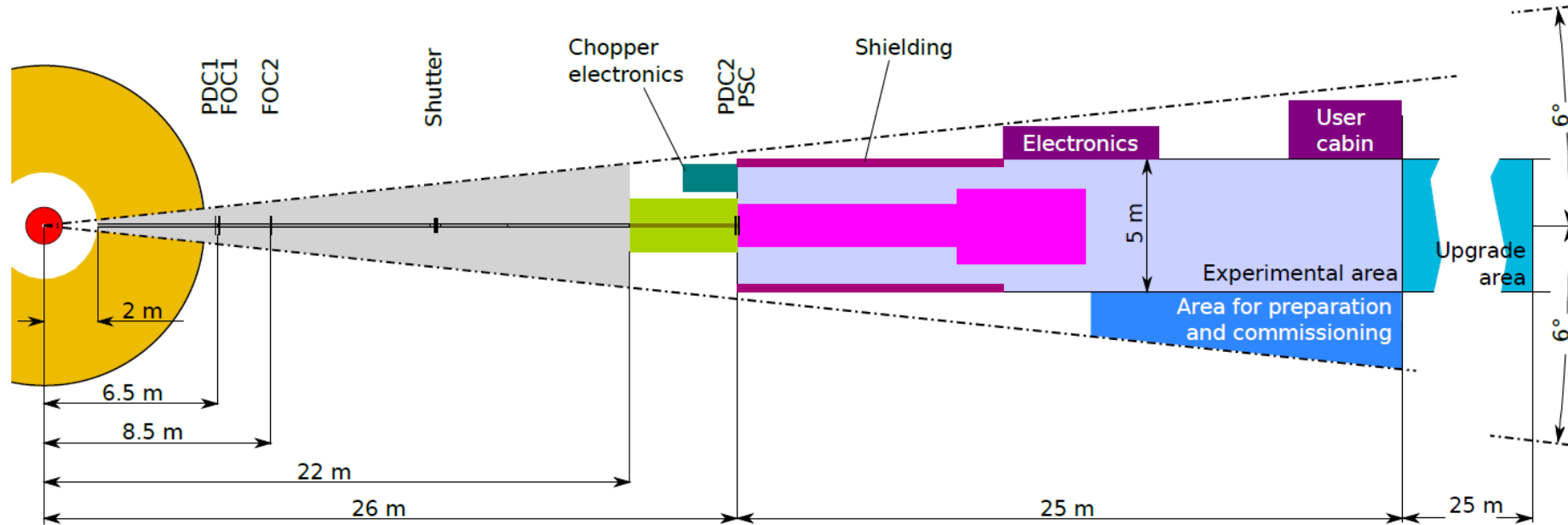
Neutrons

ESS: Beamline for Particle Physics Guide and Chopper System



Element	Dist [m]	Rad [m]	Len [m]	X-section [cm ²]	NOC	<i>m</i> value	γ^* [mrad]	λ^* [Å]
Trumpet	2.0	∞	3.6	9 × 6 → 13 × 6	1	LR: 3.5, TB: 3.0		
Straight 1	5.6	∞	0.9	13 × 6	1	3.0		
Bender 1	6.5	270	8.0	13 × 6	2	LRT: 3.0, B: 3.5	14.7	2.44
Straight 2	14.5	∞	0.4	13 × 6	2	3.0		
Bender 2 or Polarizing bender	6.5	-84	2.5	13 × 6	6	LRB: 3.0, T: 3.5	14.7	2.44
Antitrumpet	17.4	∞	4.6	13 × 6 → 11 × 7	1	LR: 3.5, TB: 3.0		

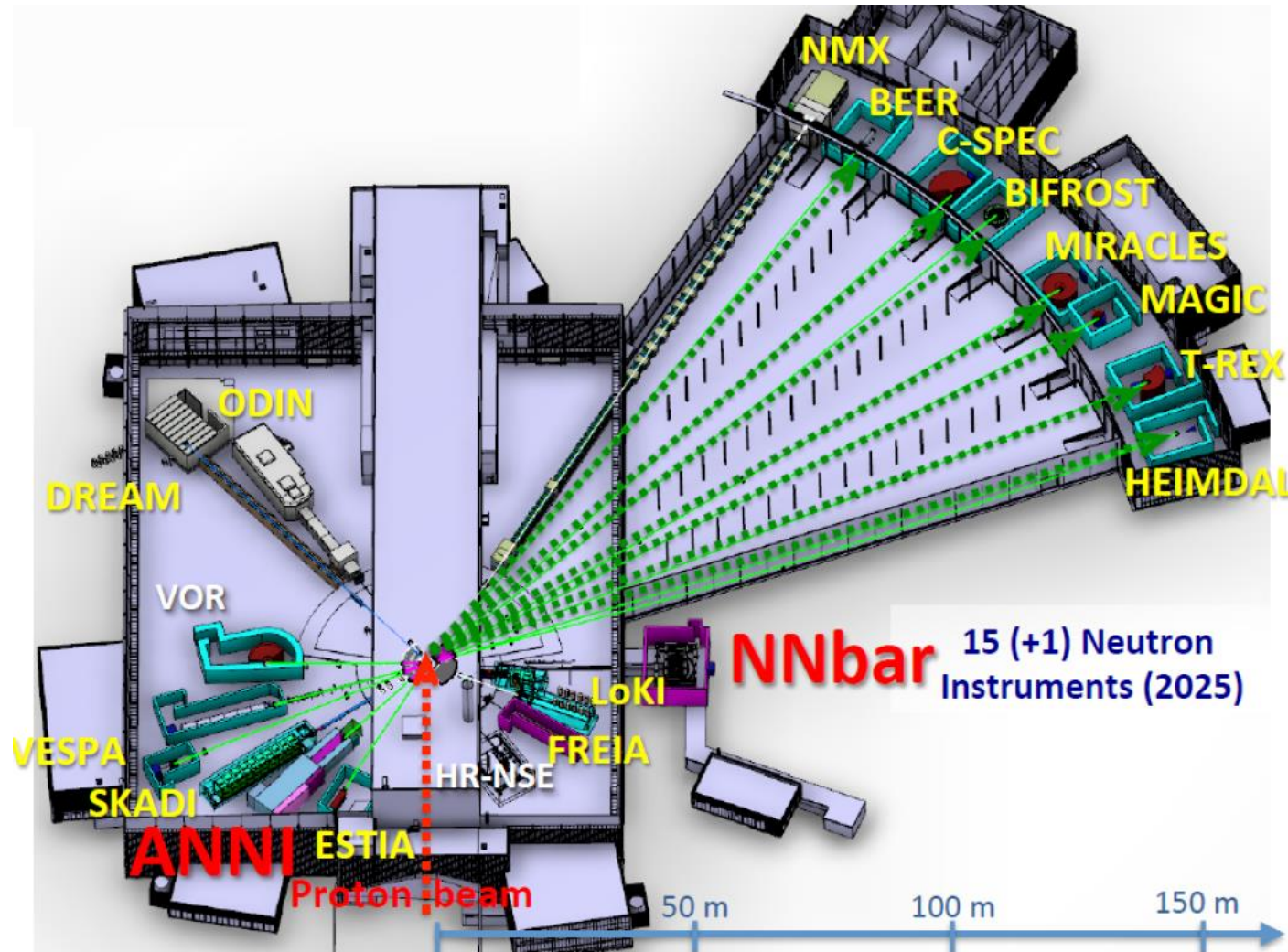
ESS: Beamline for Particle Physics Ground Floor Plan



2 frame overlap choppers (FOC1, FOC2), 2 pulse-defining choppers (PDC1, PDC2), 1 pulse-suppressing chopper PSC.

Experimental area 5 x 25 m, upgradable to 50 m length

ESS: Tentative Beam Line Location



A. Schreyer, „Workshop on Particle Physics at ESS“,
Jul. 2016, Lund, Sweden

Summary and Outlook

Pulsed cold beams enable

unprecedented statistics (or better signal / background) (at pulsed sources), and greatly improved control of systematics,

(background, edge effect, mirror effect, neutron polarisation, ...)

Newer **results on the beta asymmetry** / axial coupling constant:

consistent, small corrections, blinded*. Improvement by factor 5.

Prospects to improve neutron correlations by another **order of magnitude**

Particle Physics with Cold and Ultra-Cold Neutrons

