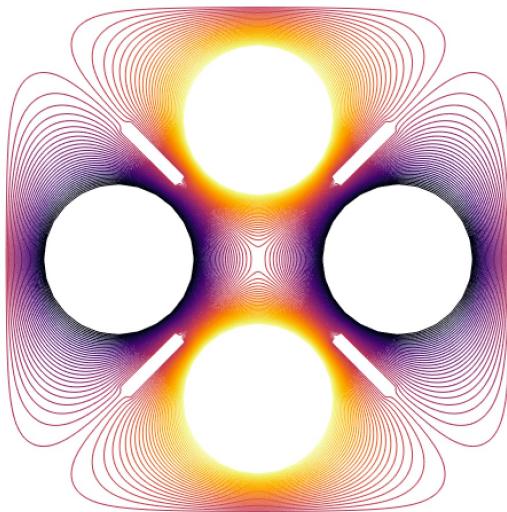
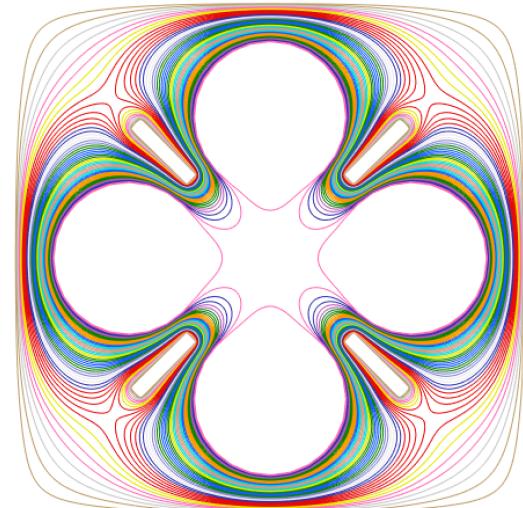


Simulating the mitigation of space charge effects with Quasi-Integrable Optics using a Paul Trap



Jake Flowerdew
Space Charge Workshop 2022
Knoxville, Tn

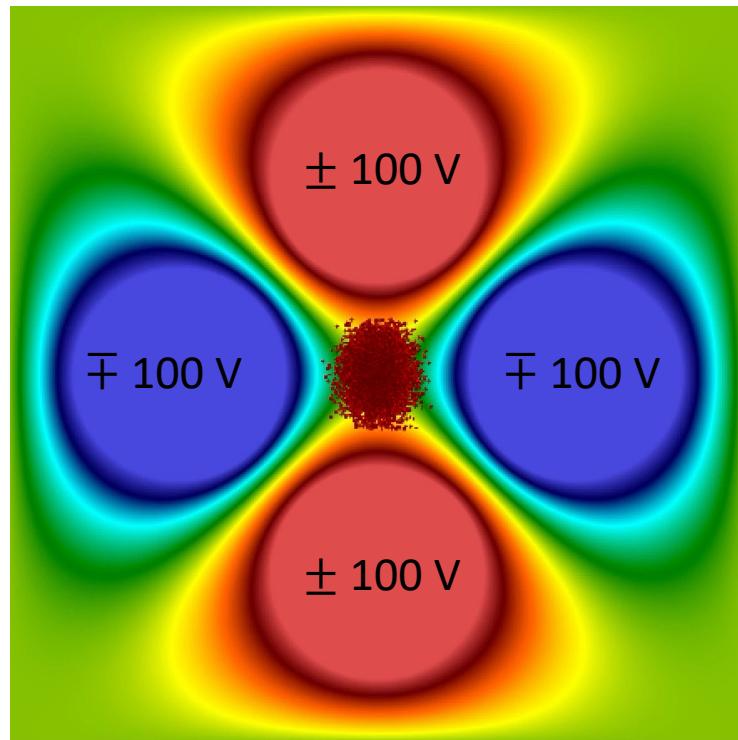


Outline

- Operation of a Paul trap
- Motivation to study accelerator physics in a Paul trap
- Quasi-Integrable Optics (QIO): Motivation to construct a nonlinear Paul trap
- Simulation results: Testing QIO with the Intense Beams Experiment (IBEX)
- Nonlinear upgrade to IBEX

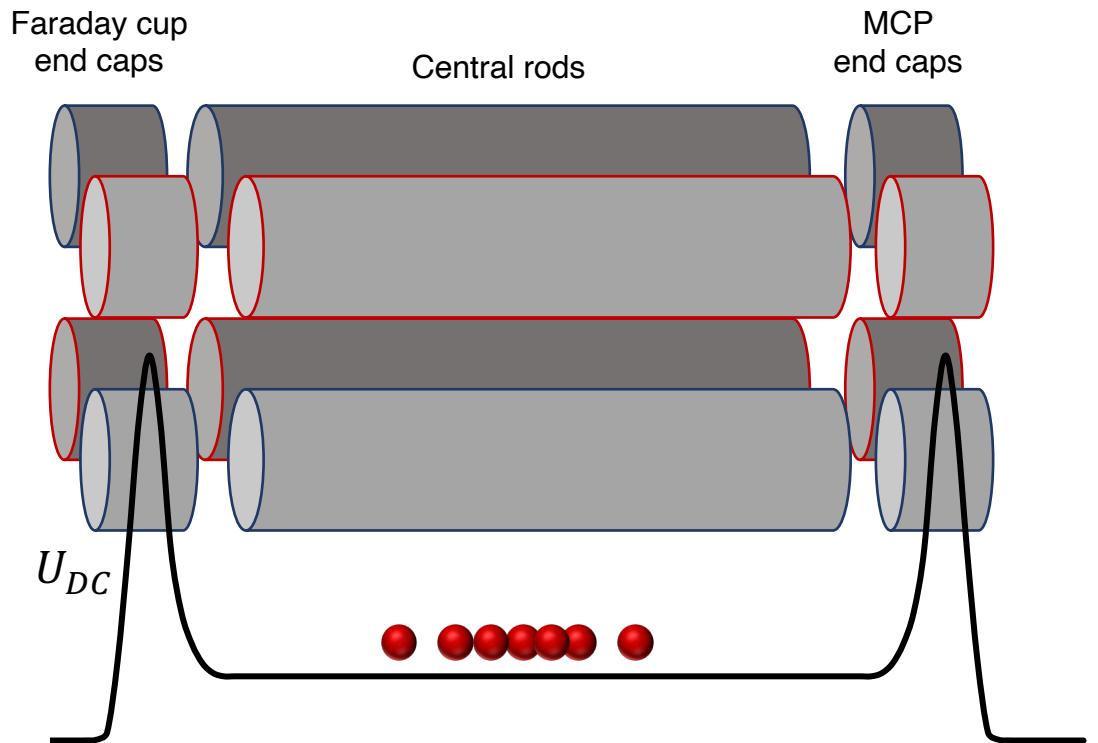
Using Paul Traps to study Accelerator physics

Transverse confinement



$$H_{\text{paul}} = \frac{(p_x^2 + p_y^2)}{2} + \frac{1}{2}K_p(\tau)(x^2 - y^2) + \frac{q}{mc^2}(\phi_{\text{sc}})$$

Longitudinal confinement



$$K_p(\tau) = \frac{2qV_Q(\tau)}{mc^2r_0^2}$$

Why study accelerator physics in a Paul trap?

- **Fast measurement times** ($1\text{ s} \sim 1,000,000$ periodic cells).
- **Large parameter space:**
 - Can easily create various different lattice types.
 - Can easily change the intensity ($\Delta Q \sim 10^{-5} - 0.1$).
- **Low energy ions** – will not damage components when lost.
- **Dispersion- and chromaticity- free** environment.
- **Cost effective** when compared to building an accelerator.



Quasi-Integrable Optics

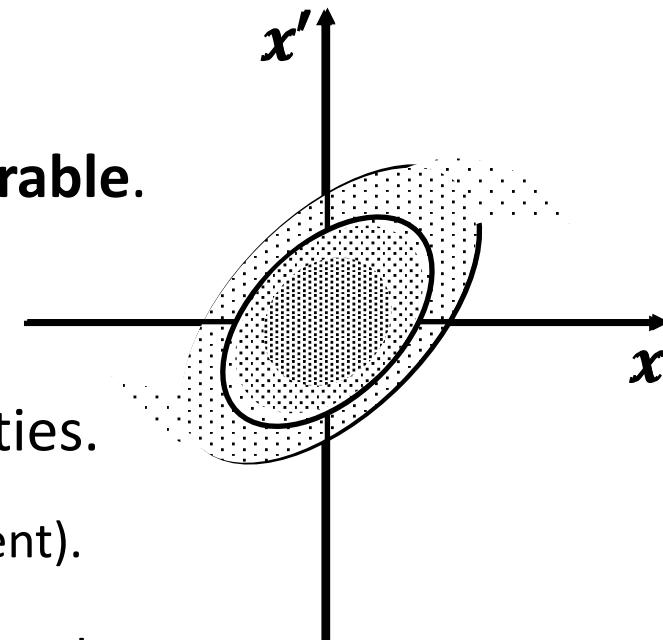
- In an ideal accelerator (linear components) the motion is **integrable**.
- Linear accelerators are susceptible to resonances.
- Use nonlinear components (i.e. octupoles) to dampen instabilities.
 - However, now the system is **non-integrable** (no longer time-independent).
- Can make a quasi-integrable system with correct octupole scaling*:

$$H_N = \frac{p_{xN}^2 + p_{yN}^2}{2} + \frac{x_N^2 + y_N^2}{2} + U(x_N, y_N),$$

$$U(x_N, y_N) = k_4 \left(\frac{x_N^4}{4} + \frac{y_N^4}{4} - \frac{3x_N^2 y_N^2}{2} \right)$$

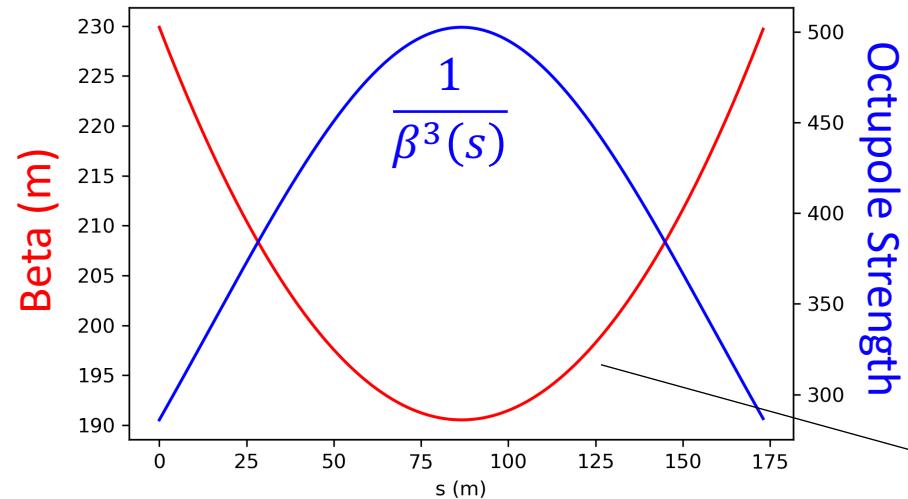
*V. Danilov and S. Nagaitsev (2010)

Bounded & regular particle motion
~~Reduced dynamic aperture~~

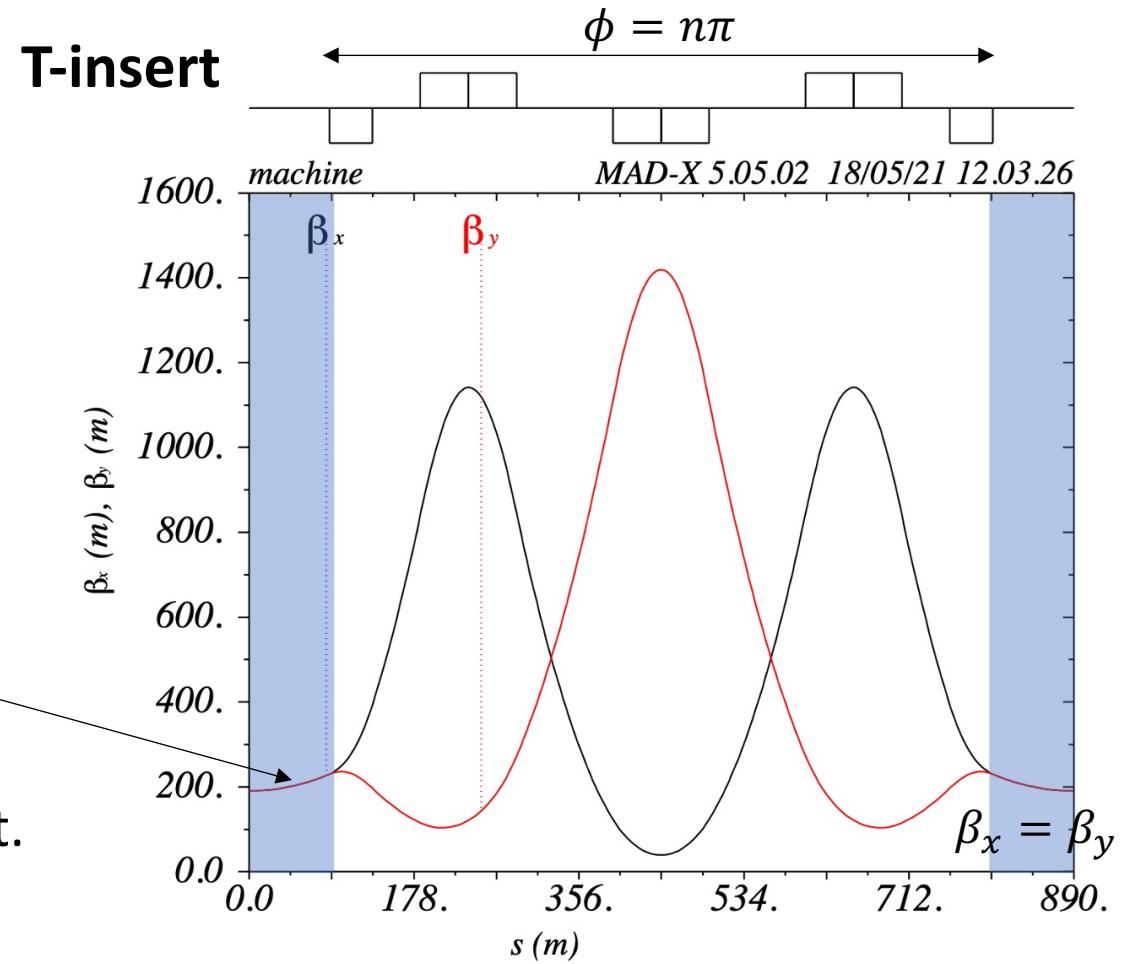


T-insert lattice for QIO

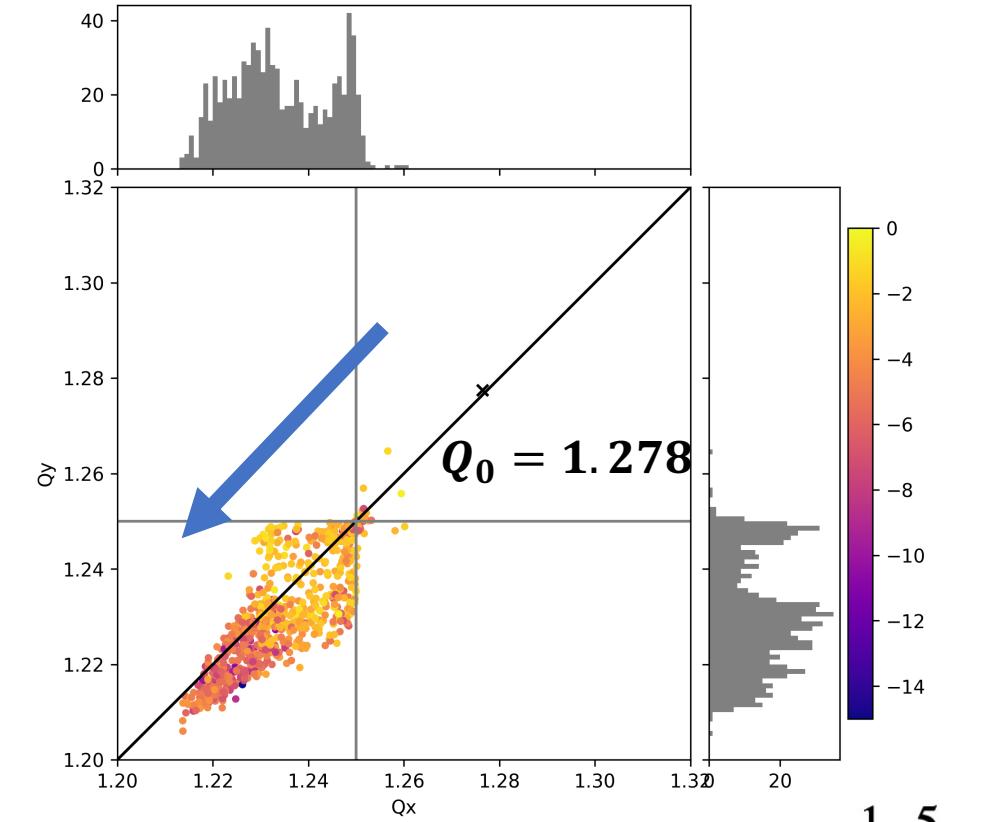
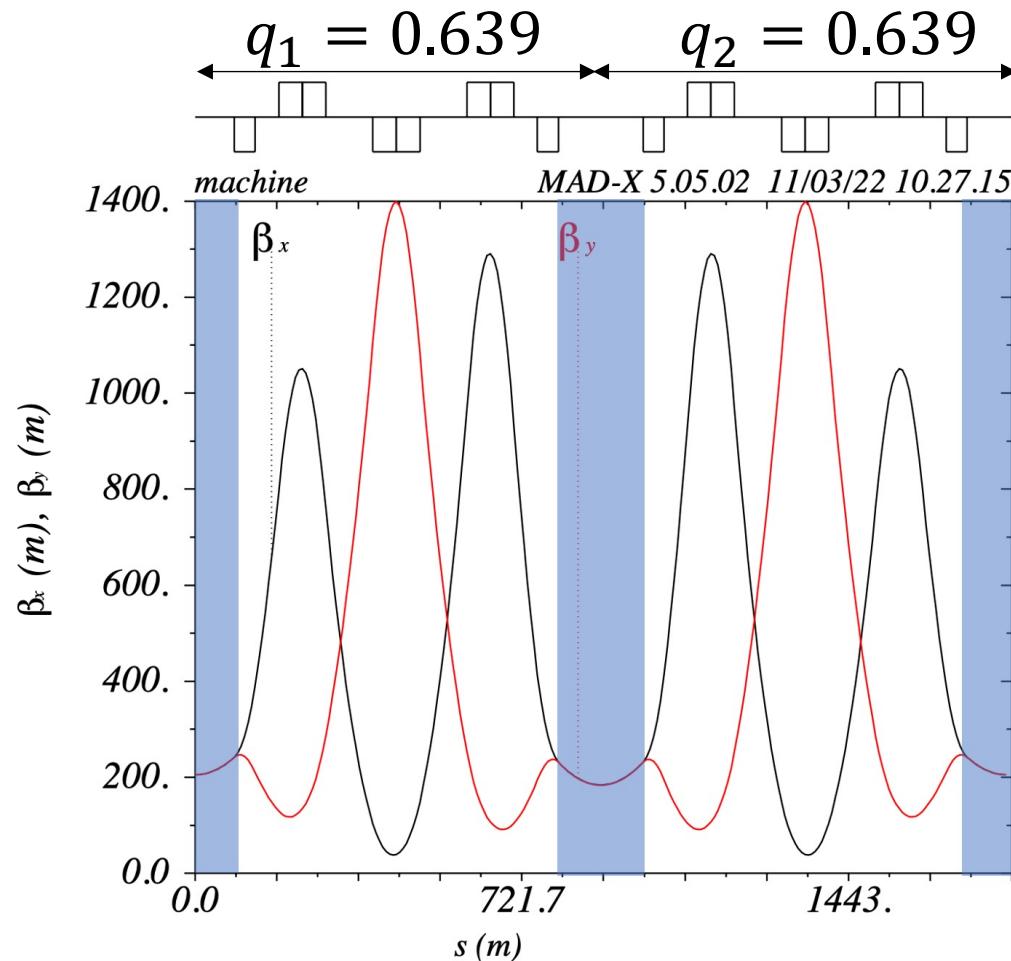
$$V(x, y, s) = \frac{k}{\beta(s)^3} \left(\frac{x^4}{4} + \frac{y^4}{4} - \frac{3x^2y^2}{2} \right)$$



- Requires $n\pi$ phase advance and $\beta_x = \beta_y$ in drift.
- $1/\beta^3(s)$ octupole scaling makes H time-independent.
- (Quasi-) Integrable lattice which is robust to small perturbations.

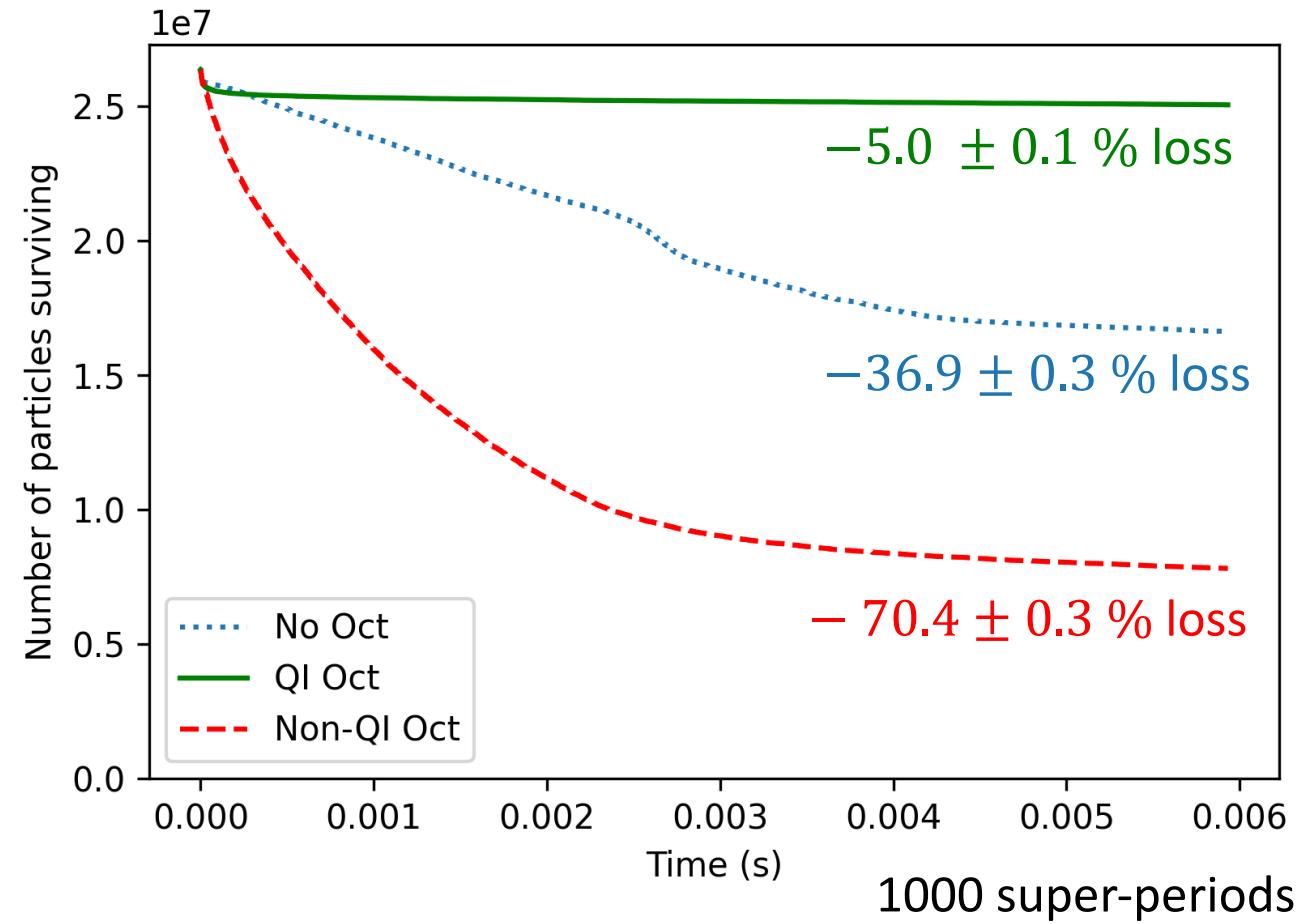
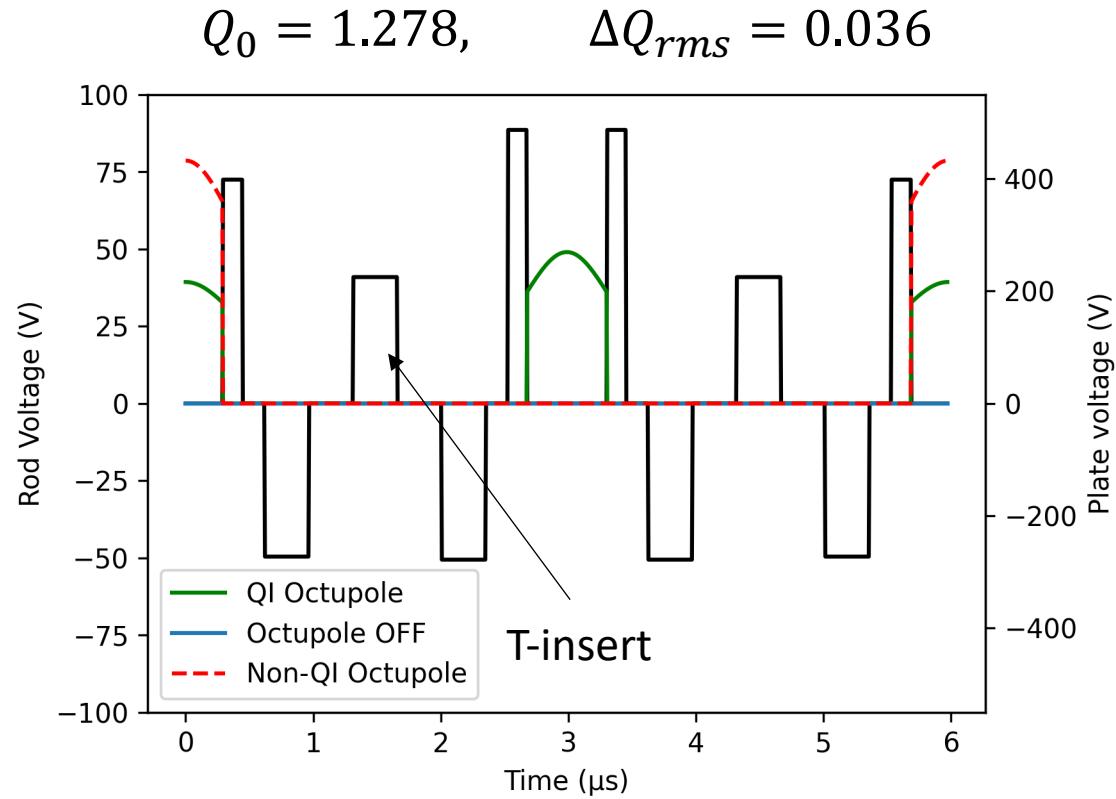


T-insert to excite coherent resonance

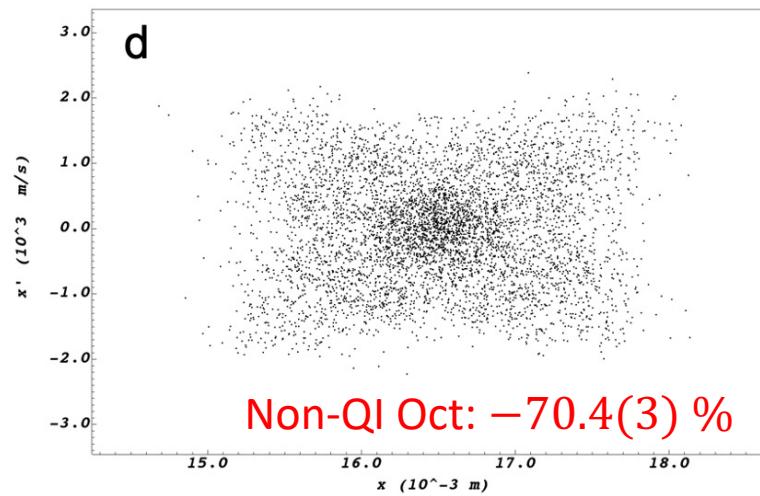
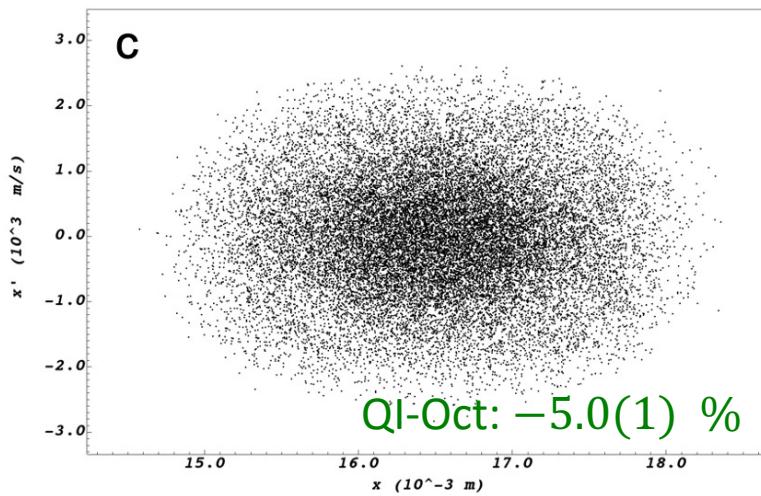
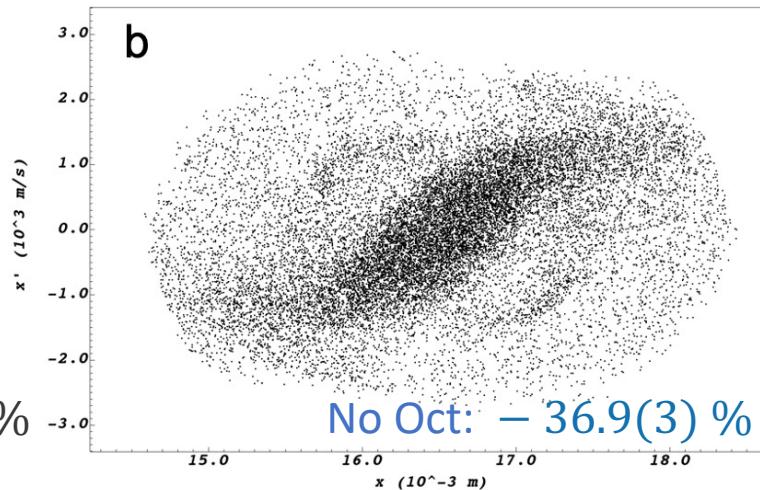
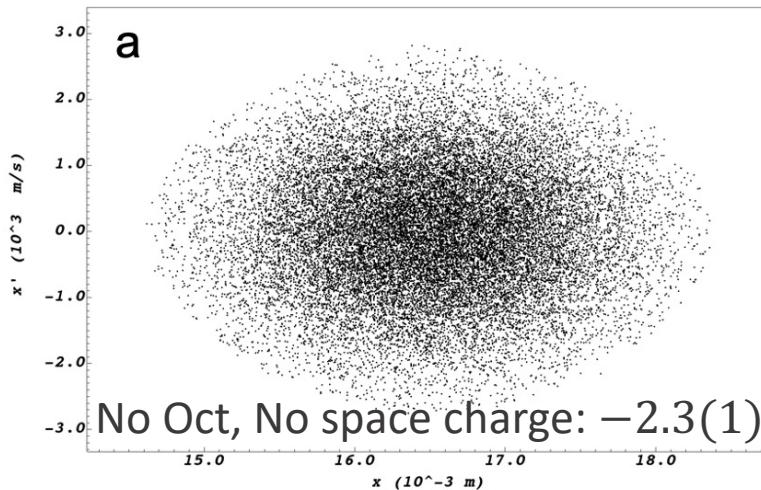


$$Q_0 + C_2 \Delta Q_{rms} = \frac{1}{2} \left(\frac{5}{2} \right)$$

VSim Numerical simulations results



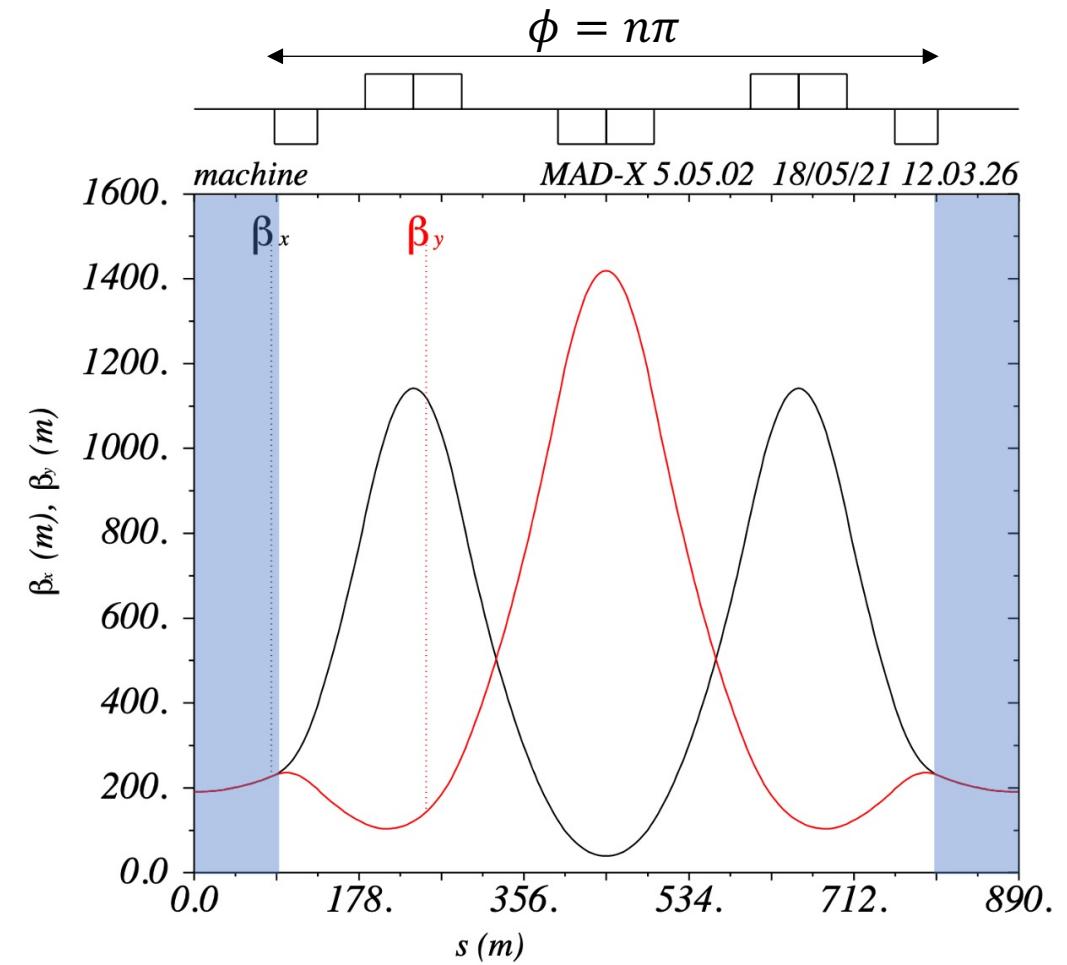
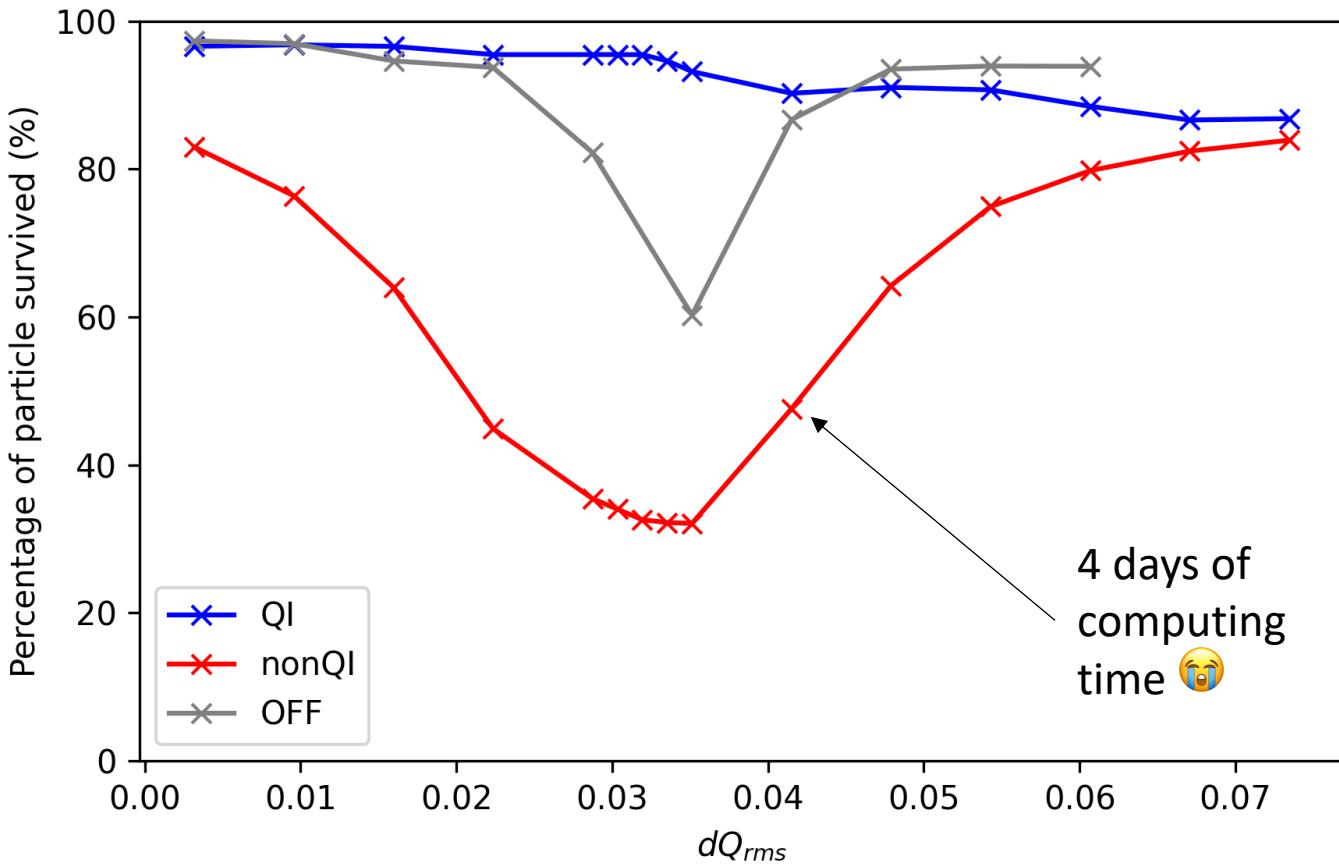
QI lattice dampens 2nd order coherent resonance



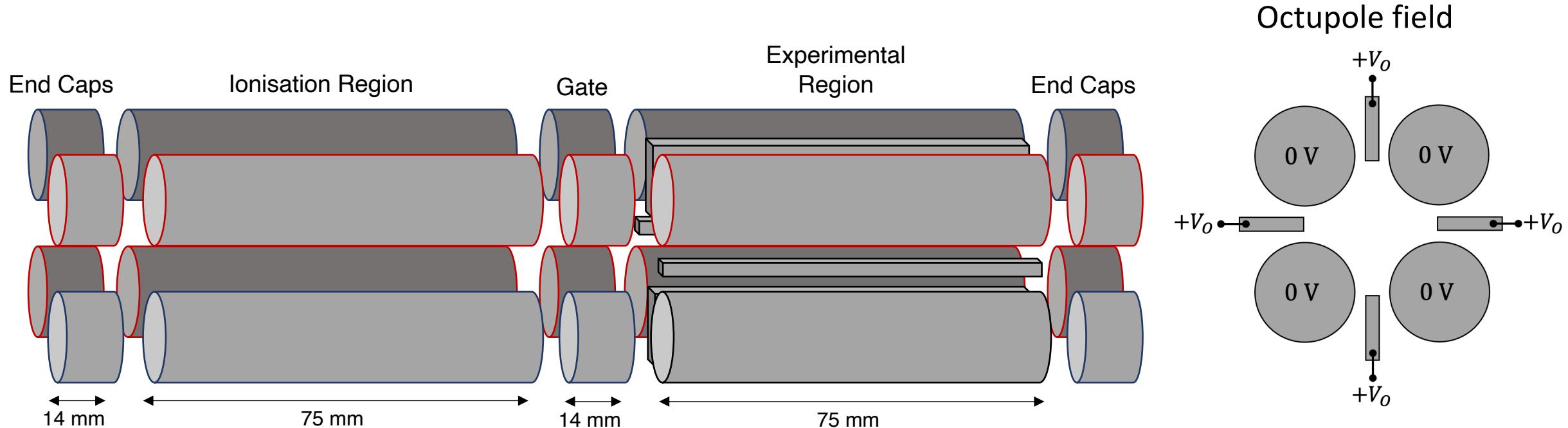
$$Q_0 + C_2 \Delta Q_{rms} = \frac{1}{2} \left(\frac{5}{2}\right)$$

$$Q_0 + \Delta Q_{rms} = \frac{5}{4}$$

Intensity scan (simulation)

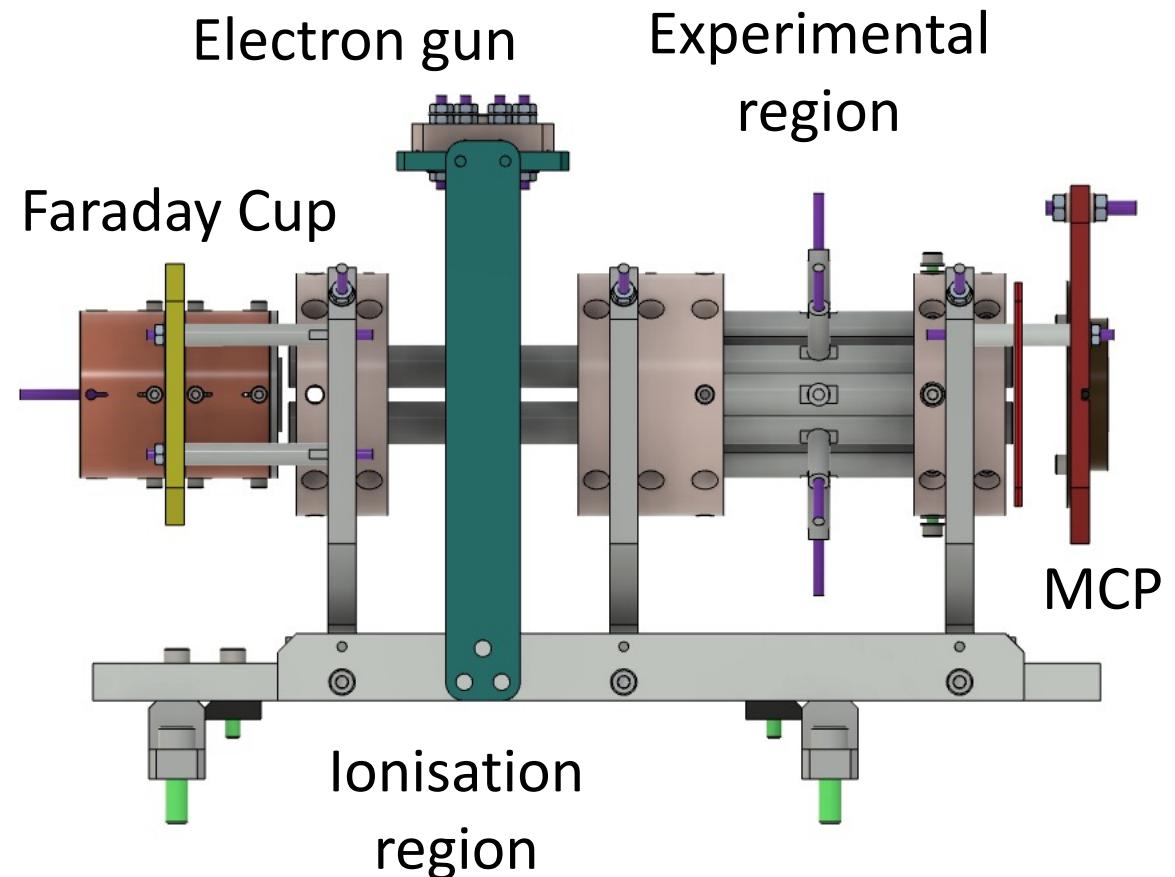
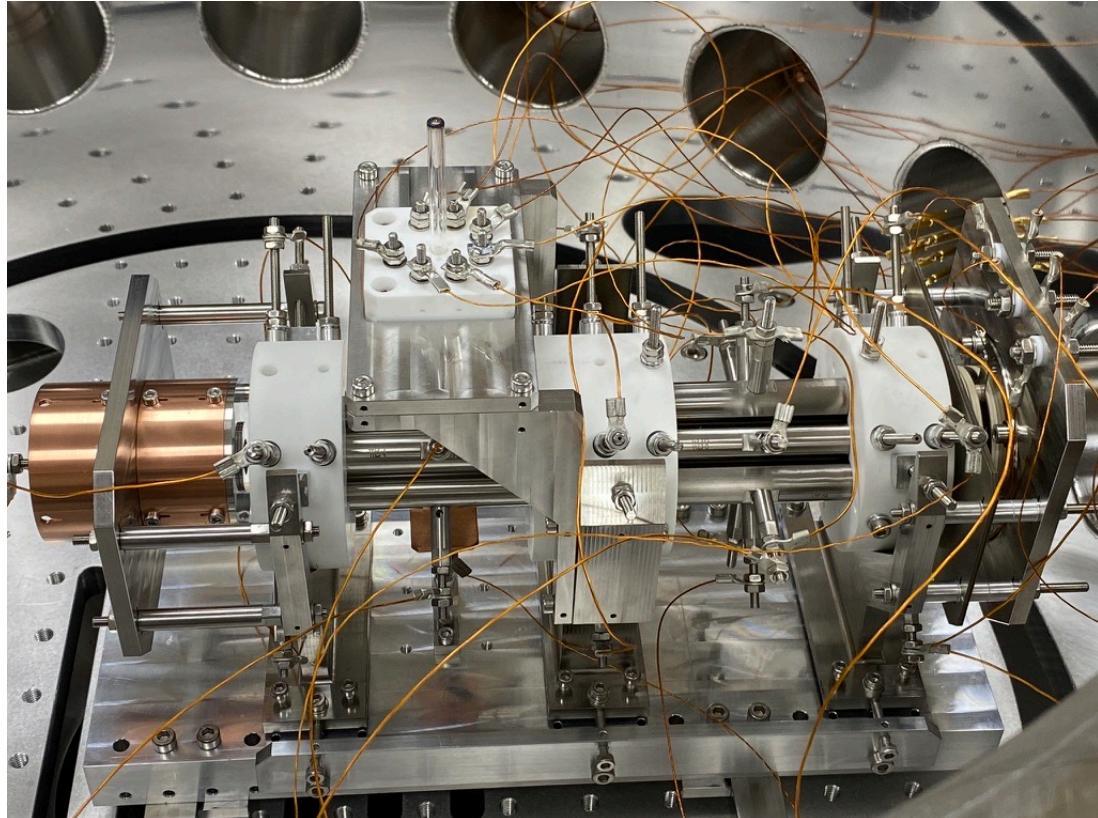


Nonlinear upgrade to IBEX



FUKUSHIMA, K., & OKAMOTO, H. (2015). Design Study of a Multipole Ion Trap for Beam Physics Applications. *Plasma and Fusion Research*, 10(0), 1401081–1401081.

IBEX 2.0!



Bridging the gap in space charge dynamics

Summarize the content of this presentation:

IBEX is a tool to test new lattice designs with space charge. In this work I presented a possible way to test the mitigation of a coherent space charge resonance with a QIO lattice in a Paul trap.

From your perspective, where is the gap regarding space charge effects?

Validating and improving upon numerical simulations of space charge.

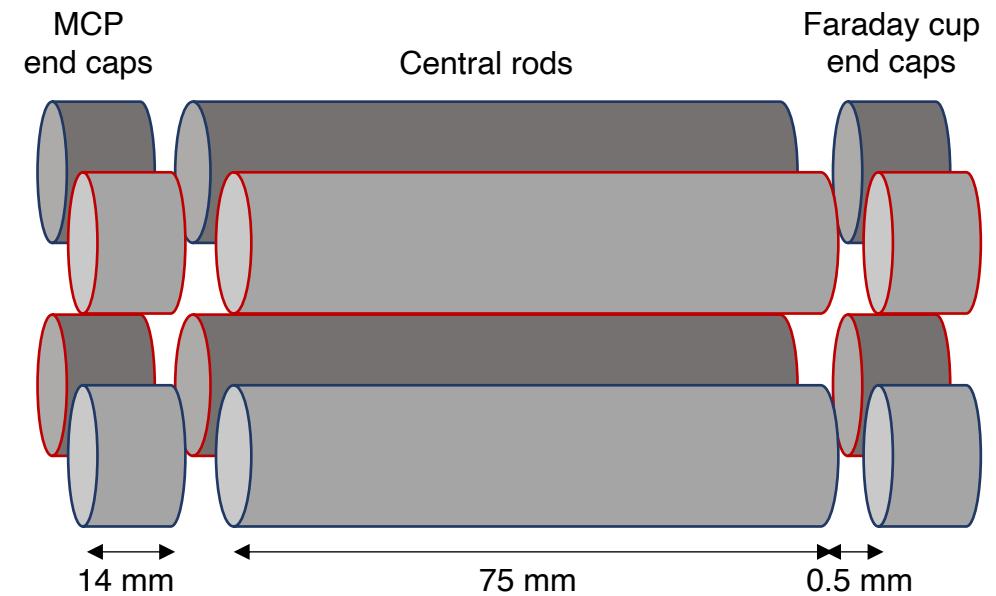
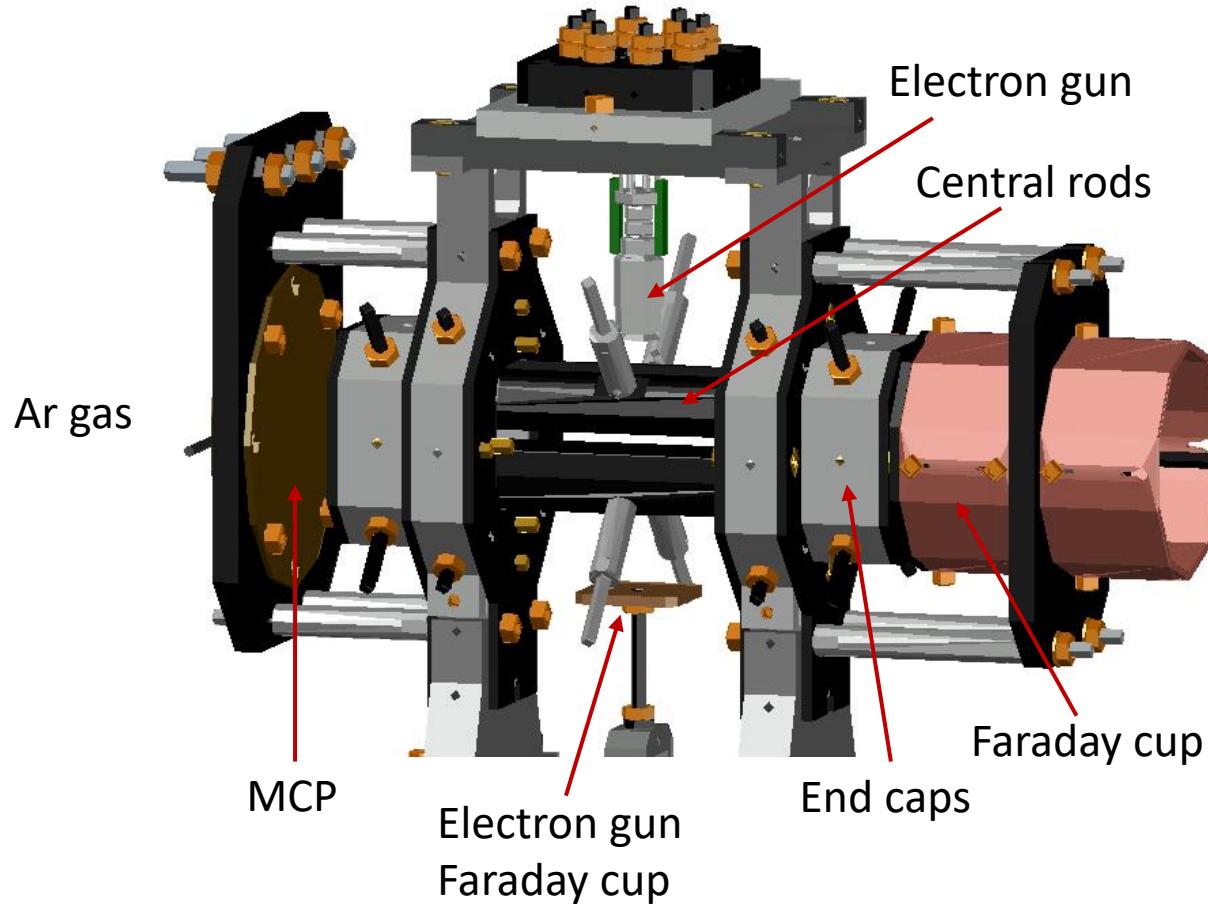
What is needed to bridge this gap?

Experimental validation of space charge effects and techniques for mitigation.

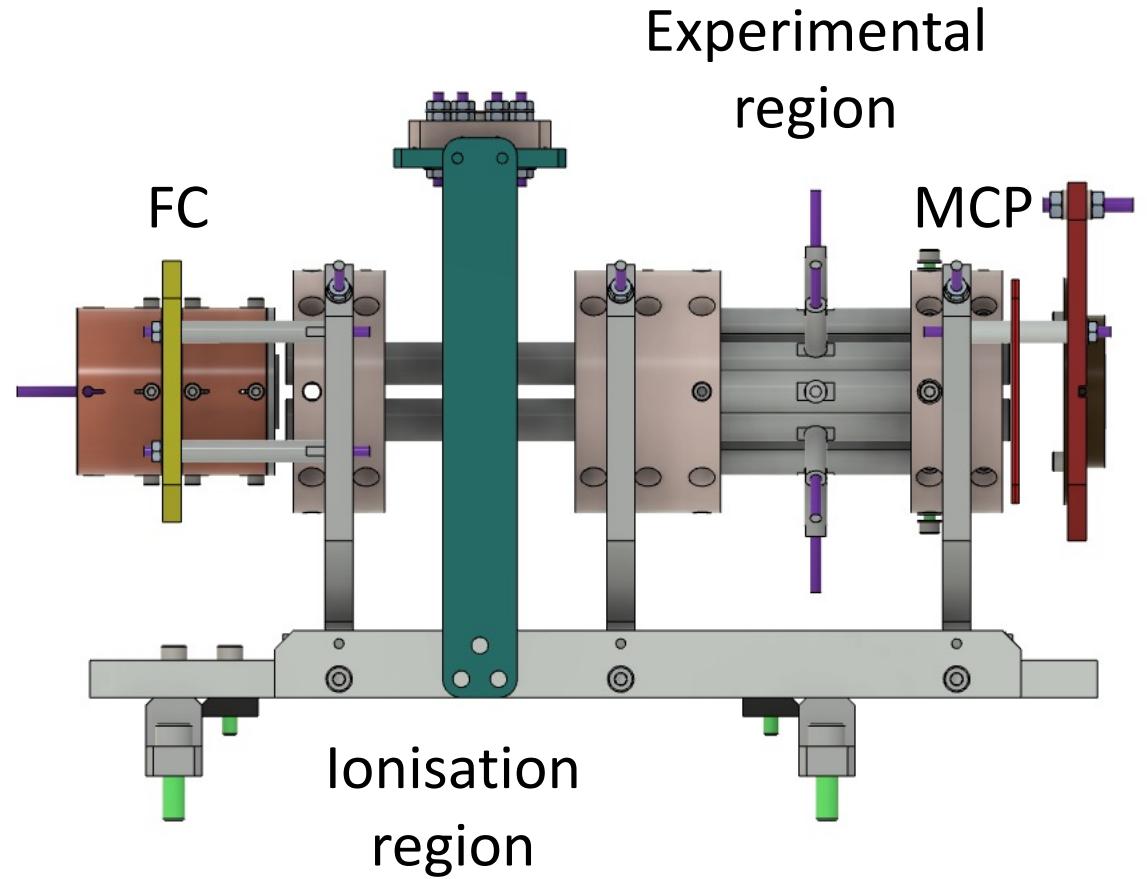
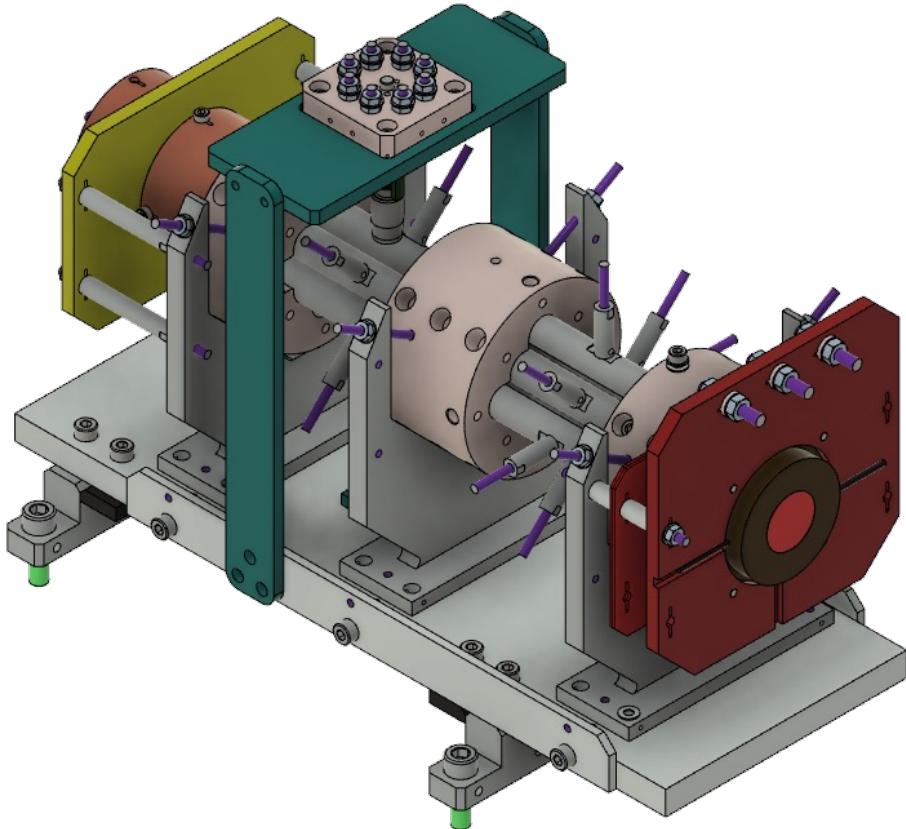
jake.flowerdew@physics.ox.ac.uk

Back up slides

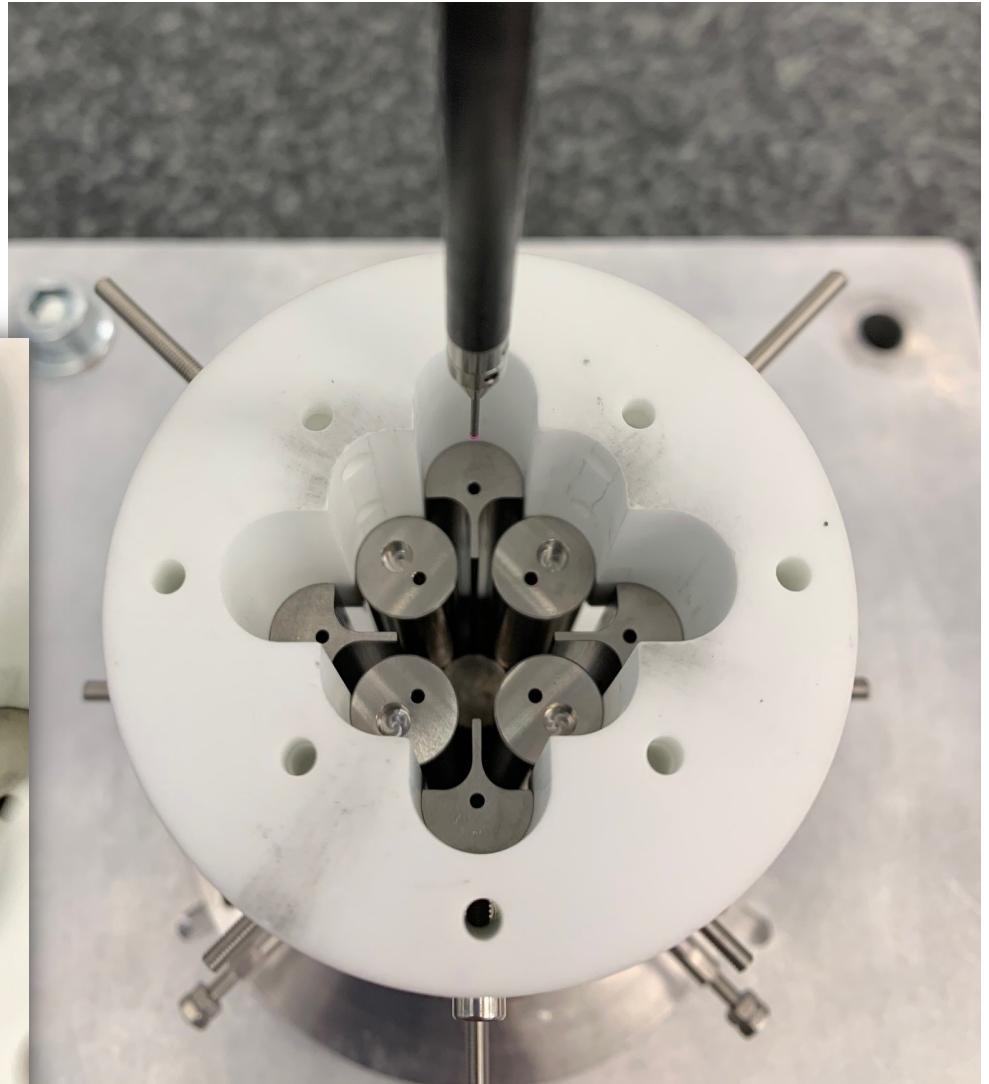
IBEX: Current Trap



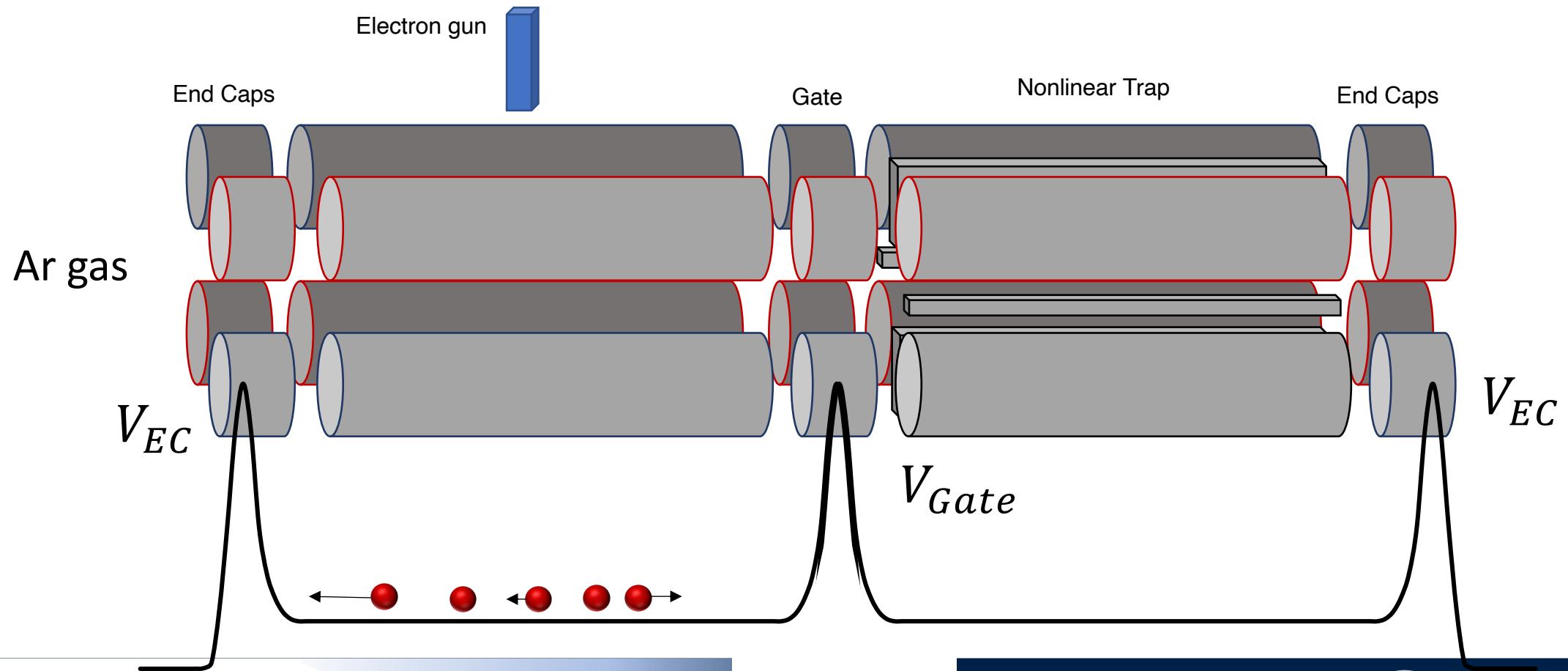
Nonlinear upgrade to IBEX



CMM measurements



Trap operation



Octupole perturbation

